



Southern Spectroscopic Project Observatory Team

Un setup remote dans l'hémisphère sud 100% automatique



Photo : © Deep Sky Chili



Rencontres du Ciel et de l'Espace 2024

par **Olivier GARDE & l'équipe 2SPOT**

Qui sommes nous ?



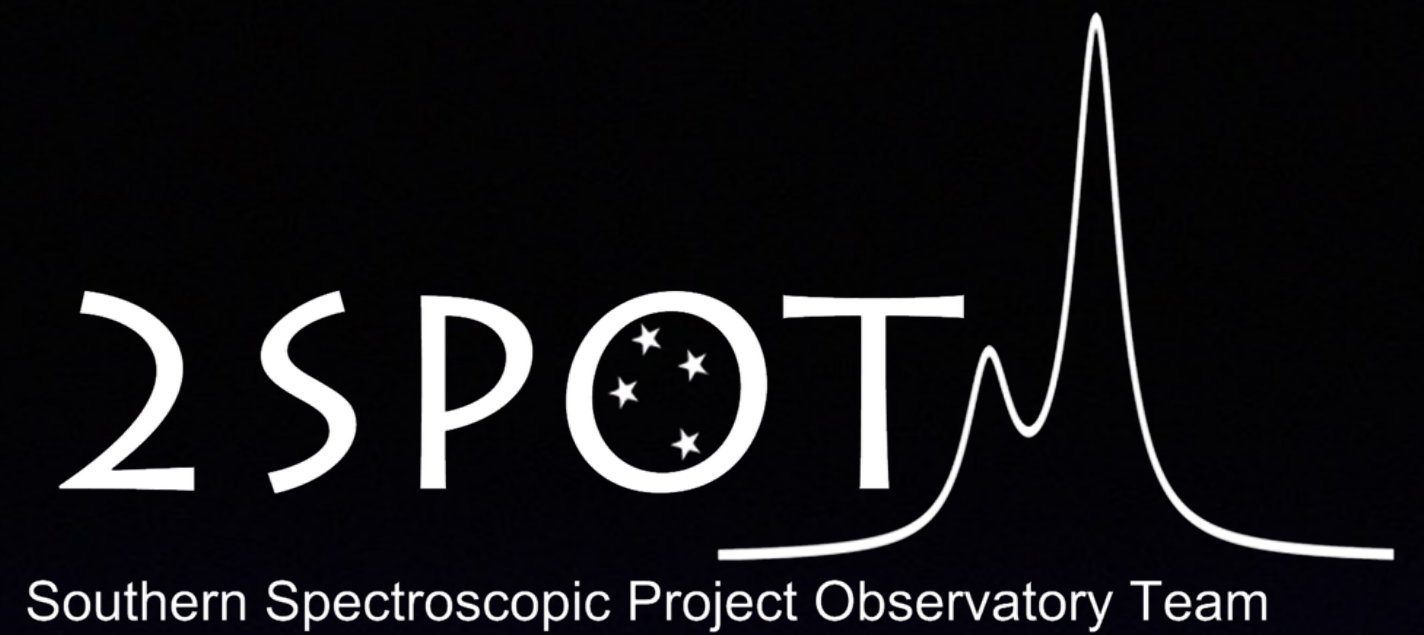
Southern Spectroscopic Project Observatory Team

- Une équipe de **5** astronomes amateurs Français
- Nous avons créé une association loi 1901 en septembre 2019
- Notre association est reconnue d'intérêt général à caractère scientifique
- Nous avons le soutien de plusieurs sociétés et instituts de recherches
- Nous participons à des projets de collaboration Pro/Am en spectroscopie



Rencontres du Ciel et de l'Espace 2024

Nos soutiens



- Plusieurs sociétés, instituts de recherche, revues scientifiques, écoles d'ingénieurs.
- Conventions avec divers observatoires professionnels
- Revendeurs et fabricants de matériel astronomiques
- Particuliers et astronomes amateurs



L'équipe



Thomas Petit

Olivier Garde

Stéphane Charbonnel

Pascal Le Dû

Lionel Mulato

Logiciels utilisés



Spec INTI

ARP





Observations **100%** automatiques

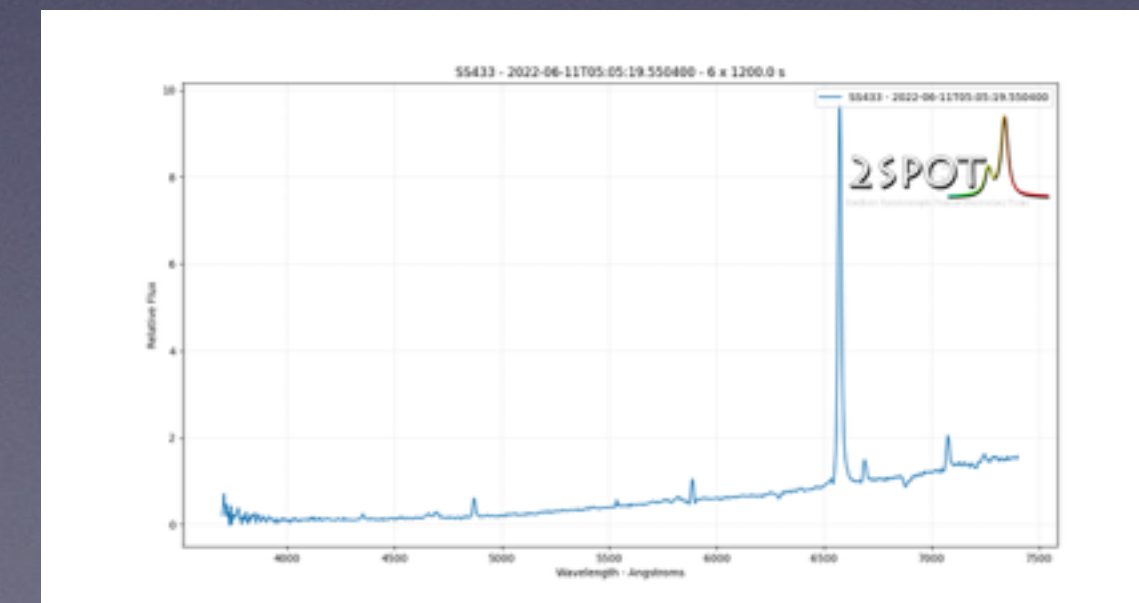
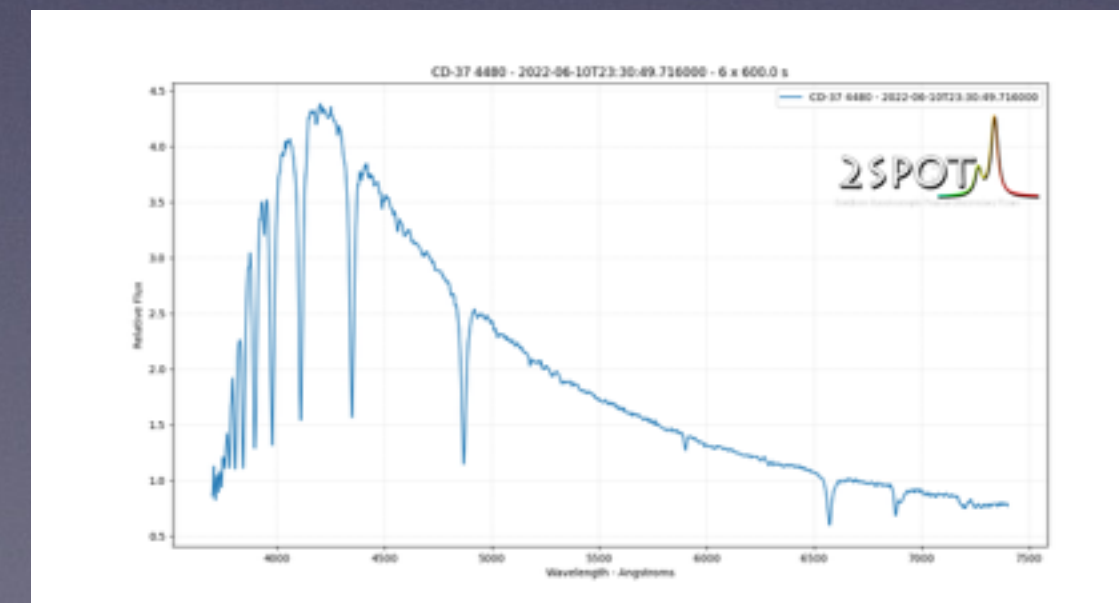
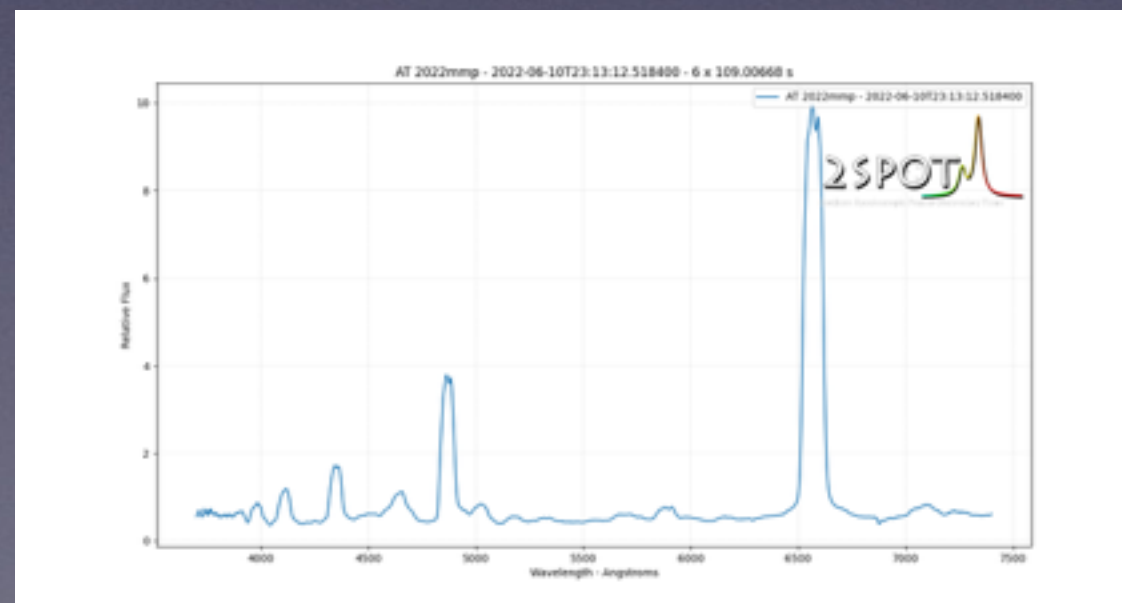
Fichier texte des divers cibles

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Objects.txt
AT_2022mnp 13 24 31.3000 -72 10 30.300 9.0 NONE NONE GUIDE NO_FOCUS Nova
CD-37__4480 08 16 12.0675 -37 47 04.222 10.89 NONE NONE GUIDE NO_FOCUS Be
U_Sco 16 22 30.7791 -17 52 43.166 10.0 NONE NONE GUIDE NO_FOCUS Nova
V644_Cen 11 43 06.5274 -60 44 04.490 10.427 NONE NONE GUIDE NO_FOCUS Be
CSI-62-12087 12 11 18.5515 -62 29 43.613 11.35 NONE NONE GUIDE NO_FOCUS Be
WRAY_15-1119 13 33 46.0353 -63 32 04.680 12.0 NONE NONE GUIDE NO_FOCUS Be
SS433 19 11 49.5647 +04 58 57.827 13.0 1200 6 GUIDE NO_FOCUS Symbiotic
HD141689 15 53 45.8361 -61 39 50.261 10.05 NONE NONE GUIDE NO_FOCUS Be
HBHA_703-05 19 12 26.9353 +06 37 44.213 11.174 NONE NONE GUIDE NO_FOCUS Be
HD_355402 20 19 21.4416 +14 54 51.455 10.87 NONE NONE GUIDE NO_FOCUS Be
```

Script Prism pour les acquisitions (**Stéphane Charbonnel**)

Process ARP (**Matthieu Le Lain** avec Spec INTI **Christian Buil**)

Résultats disponibles
1 à 2h après la fin des
observations



Dashboard (web interface)

Interface pour :

- Valider les spectres
- Générer un rapport d'observation

The screenshot displays the 2SPOT Dashboard interface. At the top, it shows the date and time in three time zones: 2022-9-30 | 09:26:36 (Chile) - 13:26:36 (UTC) - 15:26:36 (Paris). The main content area is divided into several sections:

- Dashboard Header:** Includes a "Refresh" button and a "Quick Report" button.
- Spectral Plot:** A line graph showing Flux (y-axis, ranging from -2 to 8) versus Wavelength (x-axis, ranging from 4000 to 7000). The plot shows a noisy spectrum with a prominent peak around 4000 Å and a smaller peak around 6500 Å.
- Header info:** A table of observation parameters:

Parameter	Value
OBJNAME	RX J0048 5-7302
EXPTIME2	6 x 1200.0 s
DATE-OBS	2022-09-14T04:13:00.021600
GEO_LONG	-70.853
GEO_LAT	-30.526
BSS_INST	RC12-Alpy-ATIK414Ex
OBSERVER	Olivier GARDE,L.Mulato,P.LeDu,Scharbonnel,TomPetit
JD-OBS	2459836.6757
- Tools:** A section with a button labeled "Astrometry (icrs)" and a "Get astrometry" button.
- Astronomical Images:** Two grayscale images of the field of view. The top image is a zoomed-in view of the star, and the bottom image is a wider field of view. Both images have a "Lum." (Luminosity) slider on the right and a "Cont." (Contrast) slider at the bottom.



Choix du site au Chili

DeepSkyChile
RENT CHILEAN SKY

- **320** nuits de ciel dégagé par an
- 1700m d'altitude
- Fond de ciel noir à Mag. 21-22
- Seeing très souvent $< 1''$

Lat : $30^{\circ} 31' S$

Lon : $70^{\circ} 51' W$

www.deepskychile.com



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Photo : © Google Map

Nos voisins au Chili

Cerro Tololo

Cerro Pachon



Photo : © Deep Sky Chili



Photo : © Google Map



Cerro Tololo Inter-American Observatory



Large Synoptic Survey Telescope



Southern Astrophysical Research Telescope



Gemini South



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L'observatoire Deep Sky Chile

- 5 hangars collectifs
- Panneaux solaires
- Internet avec de la 4G
- Ouverture/fermeture automatique
- Pas de pollution lumineuse



Photo : © Deep Sky Chile

- Chaque hangar abrite plusieurs télescopes



Photo : © Deep Sky Chile



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Notre 1er setup

- Télescope Ritchey-Chretien 305mm
- Monture GM3000 HPS 10Micron
- Spectrographe ALPY 600 R=600
- ATIK 414 Ex (camera spectrale)
- ATIK 314L+ (autoguidage)
- Chercheur EV Bony 60mm
- ASI 178 MM (pour le chercheur)



Chercheur électronique



Spectrographe Alpy 600



Test de notre setup à l'OHP en 2020



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Armoire électrique du 1er setup

Alimentation 24V pour la monture

Relais

Alimentation 12V pour les CCD
et autres périphériques

IPX 800 (interrupteurs sous IP)

Switch Ethernet

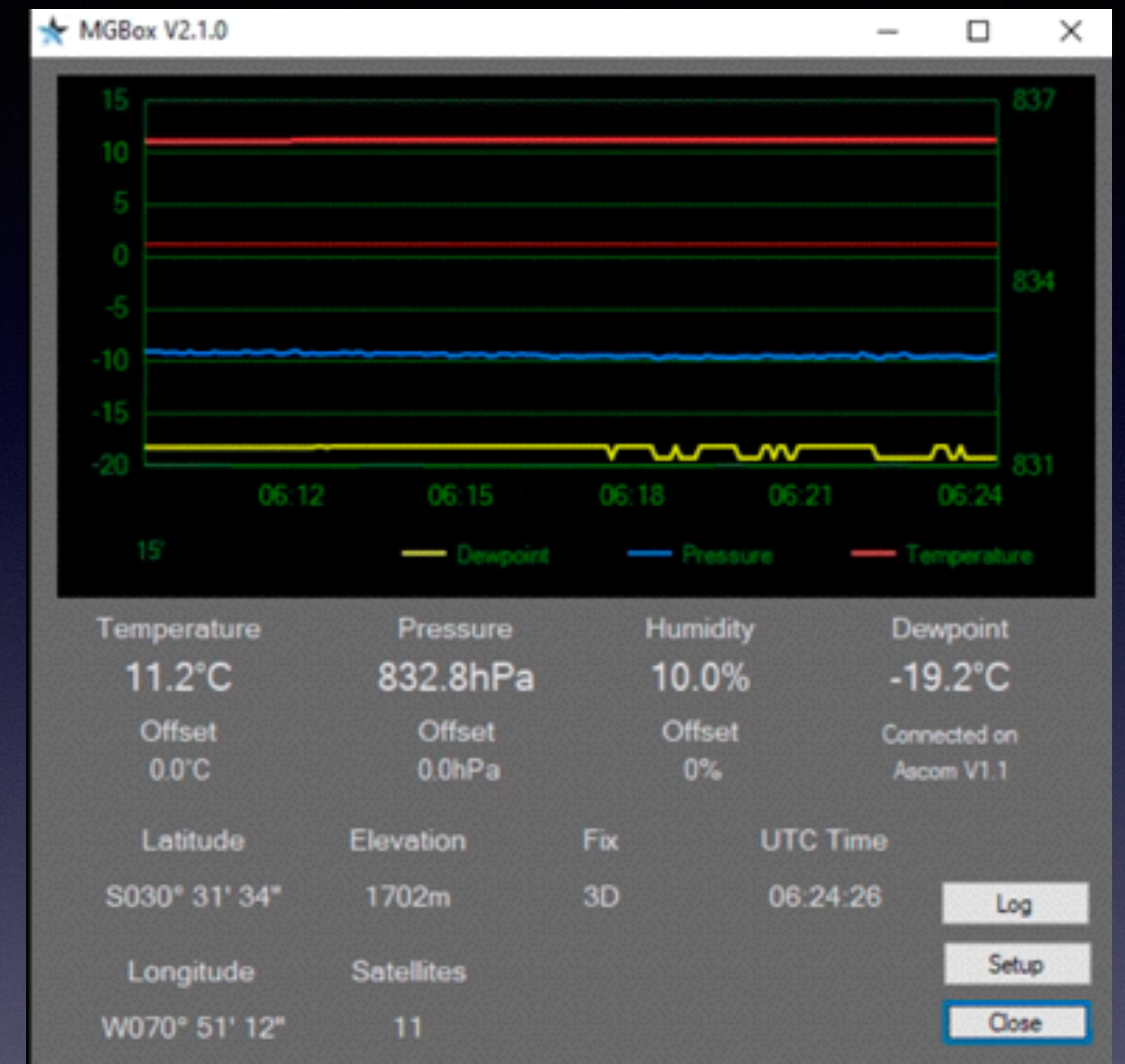
PC local et écran



Contrôle à distance des alimentations



Interface web
IPX 800



Pression & Temperature
MG BOX V2



Expédition du matériel

Départ : France
15 mars 2021

- 2 caisses en bois
- Poids: 334 Kg
- Expédition par bateau
- 45 jours de voyage
- 15000 km parcourus

Arrivé : Chile
1er mai 2021

Vidéo du voyage

www.youtube.com/watch?v=SGwaHtYerNY



Photo : © Deep Sky Chile



Rencontres du Ciel et de l'Espace 2024

Notre programme d'observation

- Confirmation de candidates NP
- Etoiles Be
- Etoiles symbiotiques et cataclysmiques
- Confirmation d'étoiles symbiotiques
- Novae
- Supernovae
- Comètes
- Flickering (WR 11 et autres cibles)
- Autres évènement dans le ciel ?

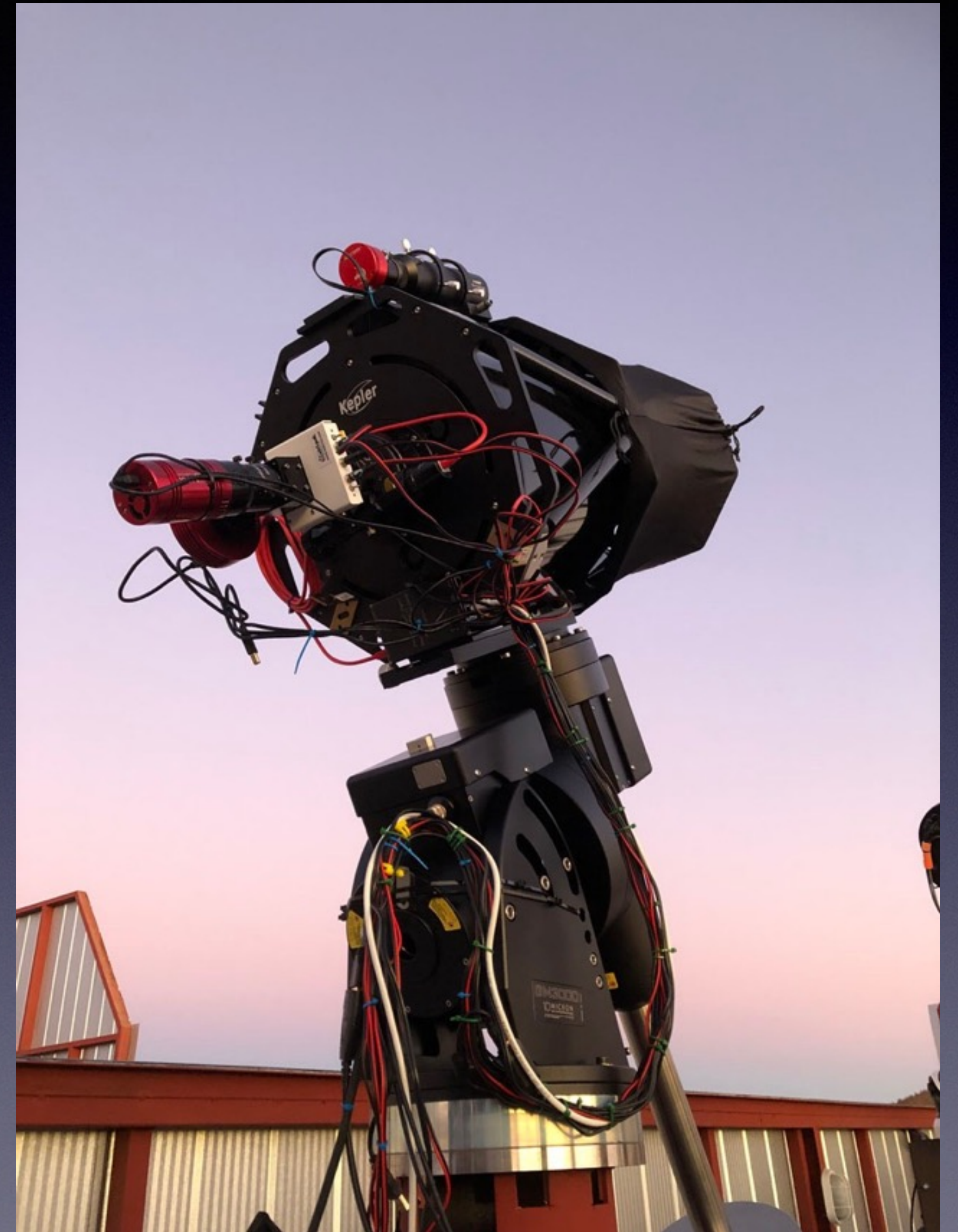
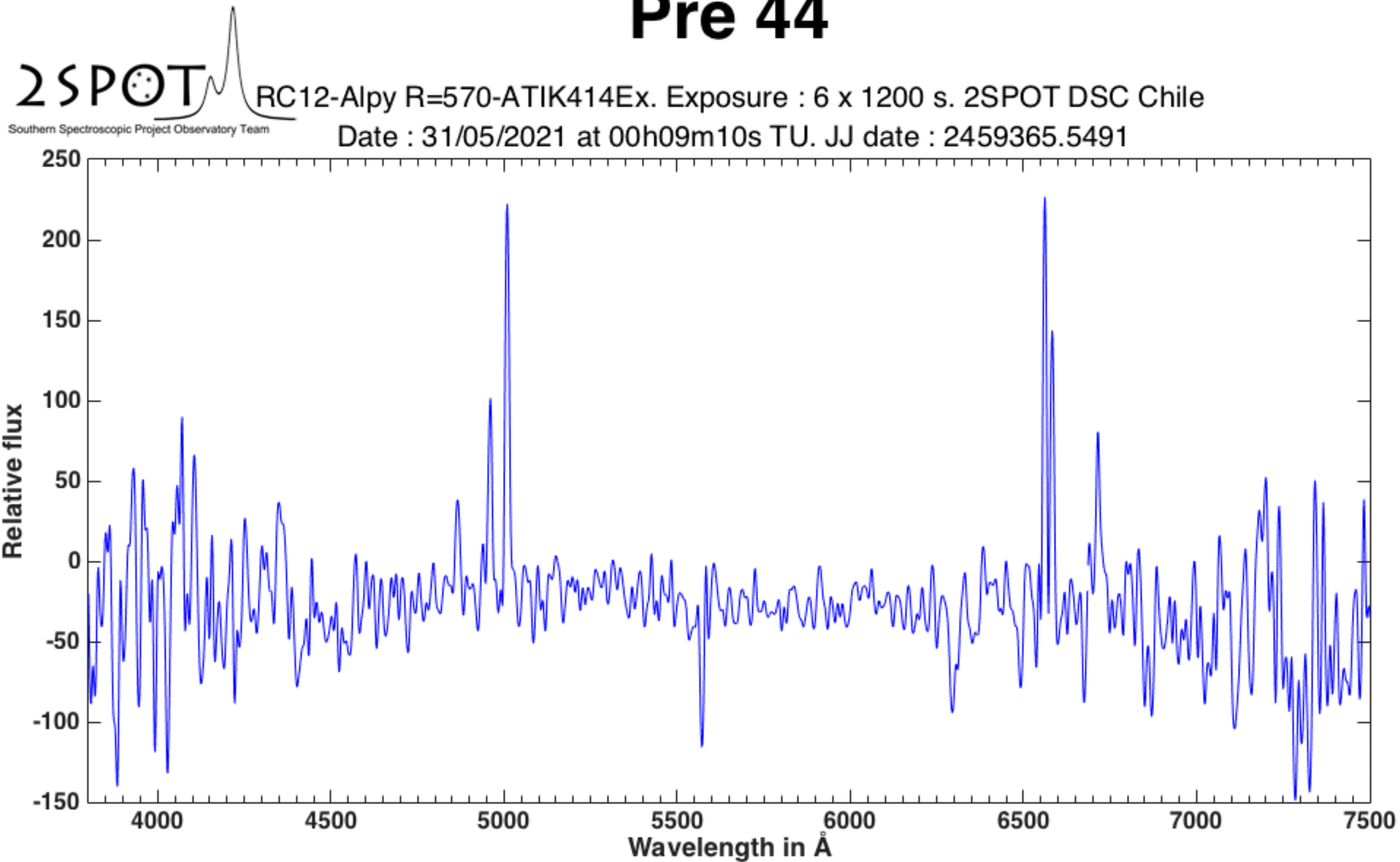


Photo : © Deep Sky Chile



Quelques résultats

Pre 44



NP candidate Pre 44 : une cible très faible



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Quelques résultats : la Sn 2021 pit

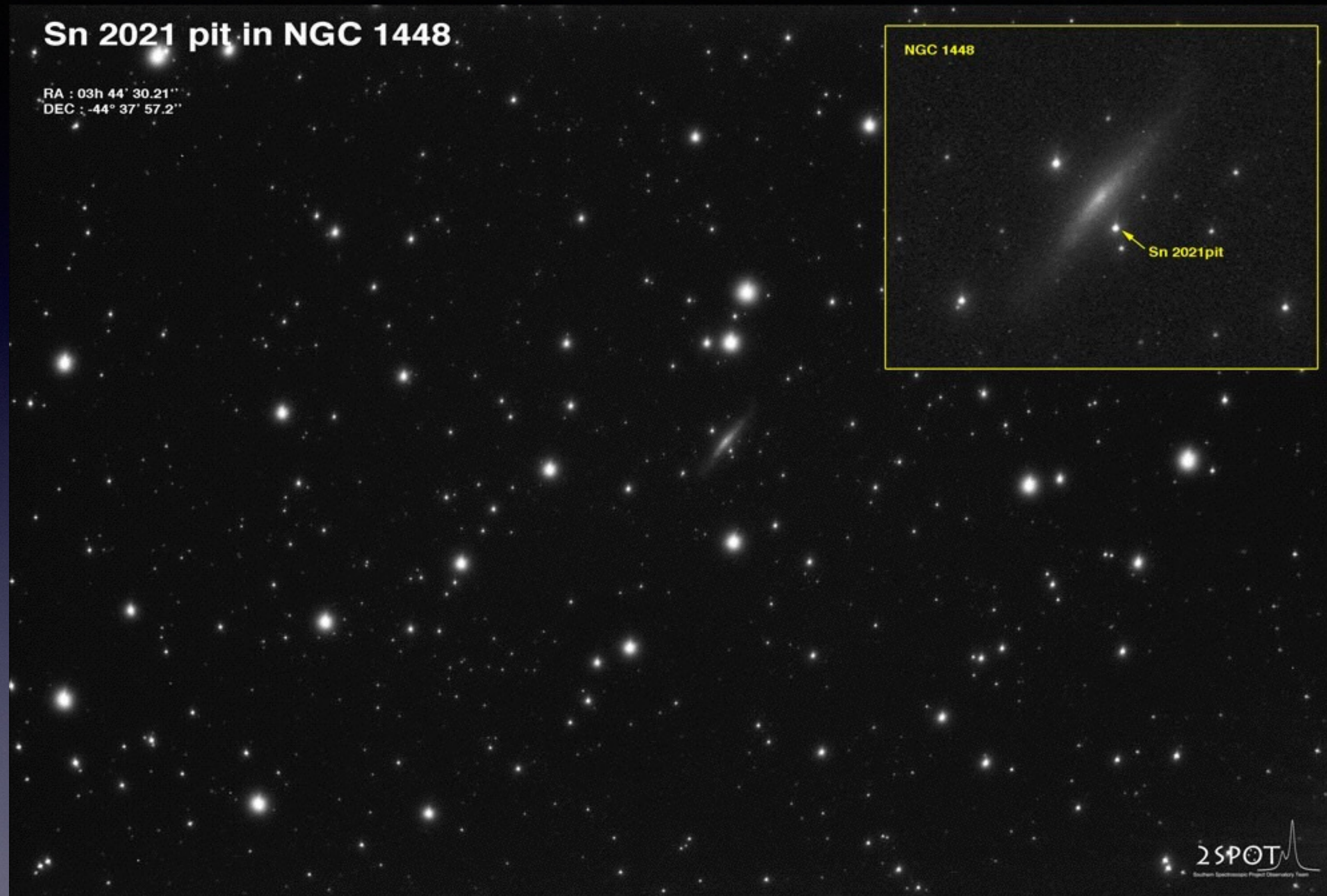
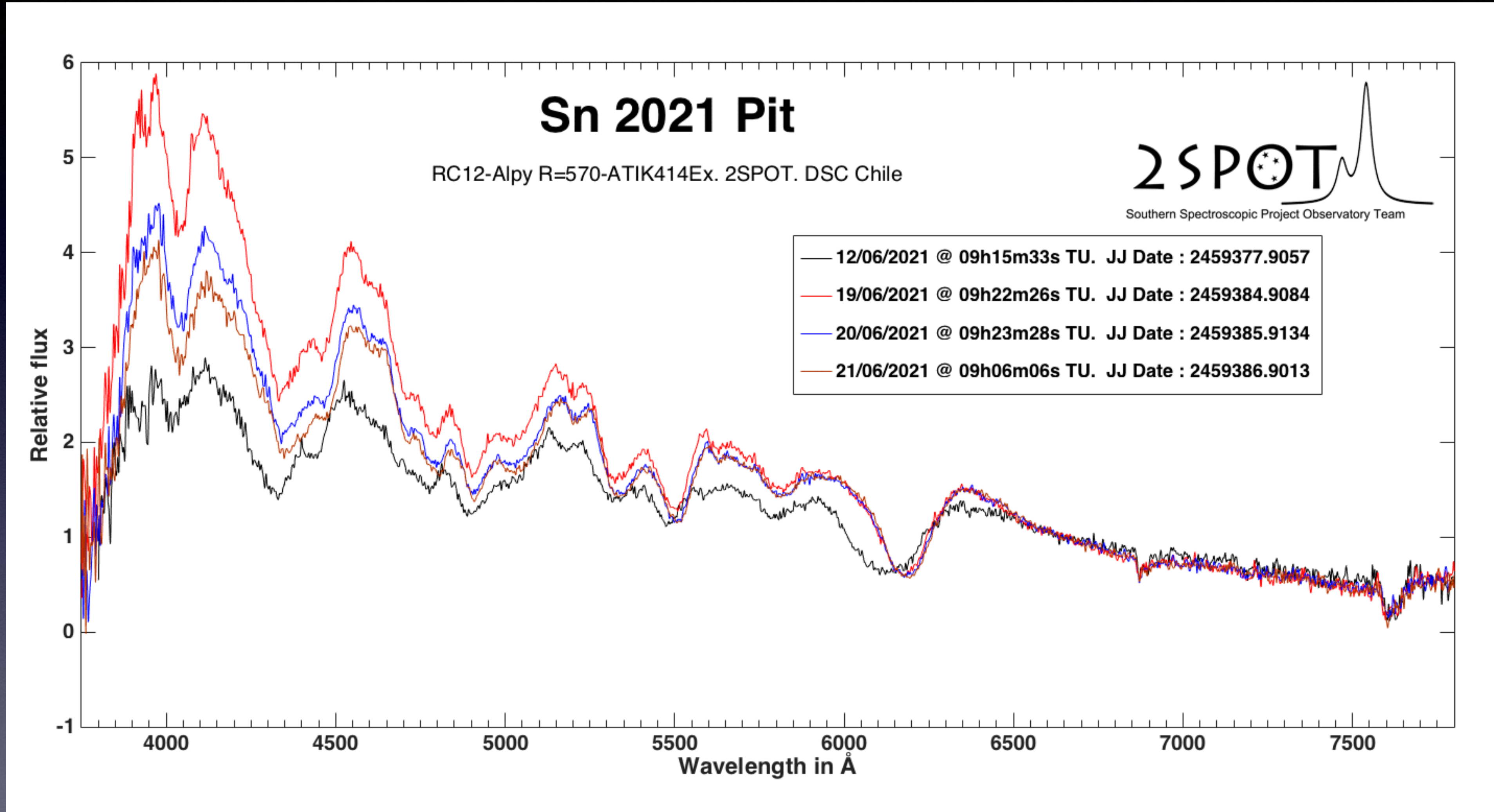


Image réalisé avec notre chercheur électronique

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Quelques résultats : la Sn 2021 pit

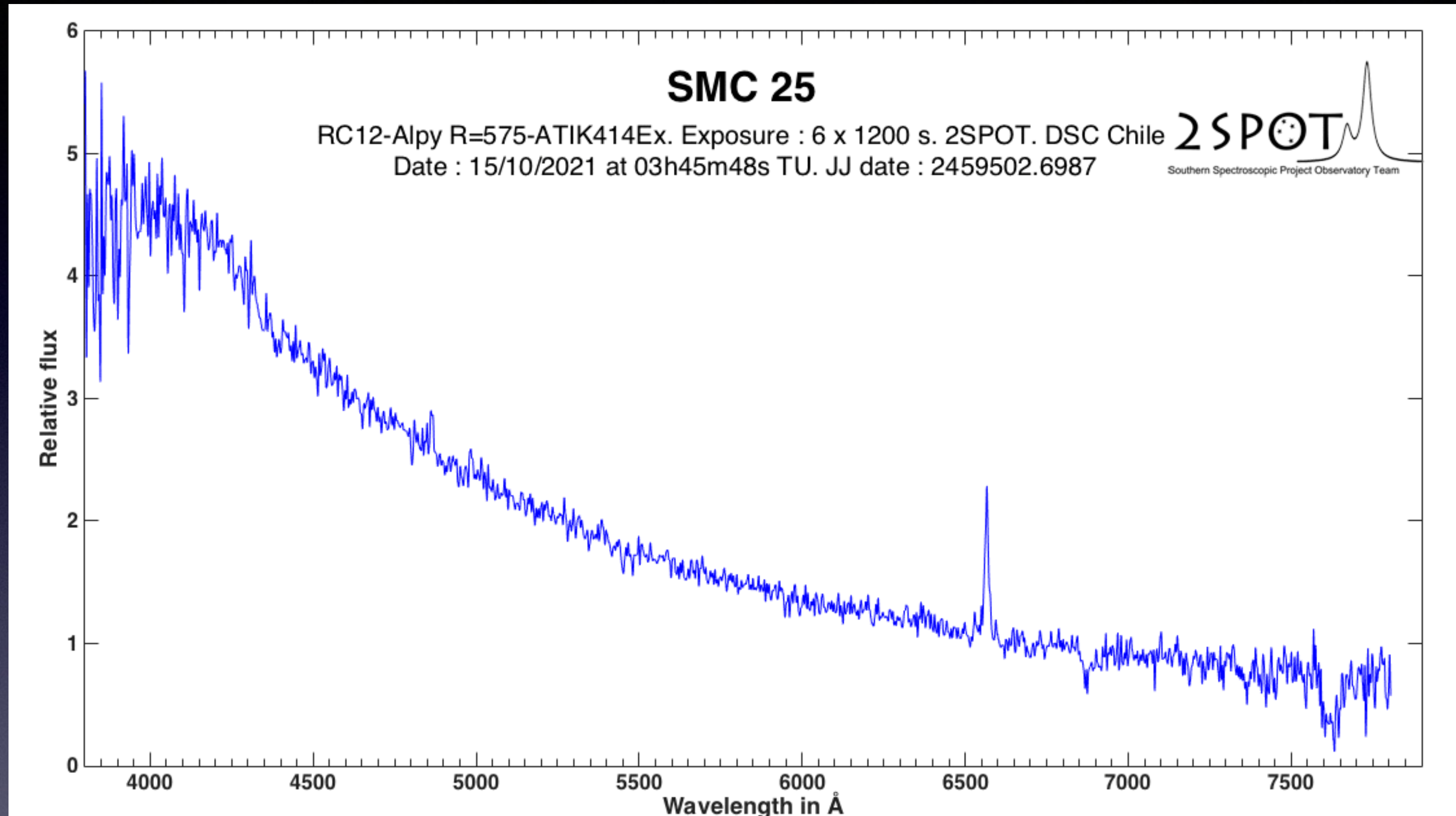


Evolution du spectre sur 4 nuits

Rencontres du Ciel et de l'Espace 2024

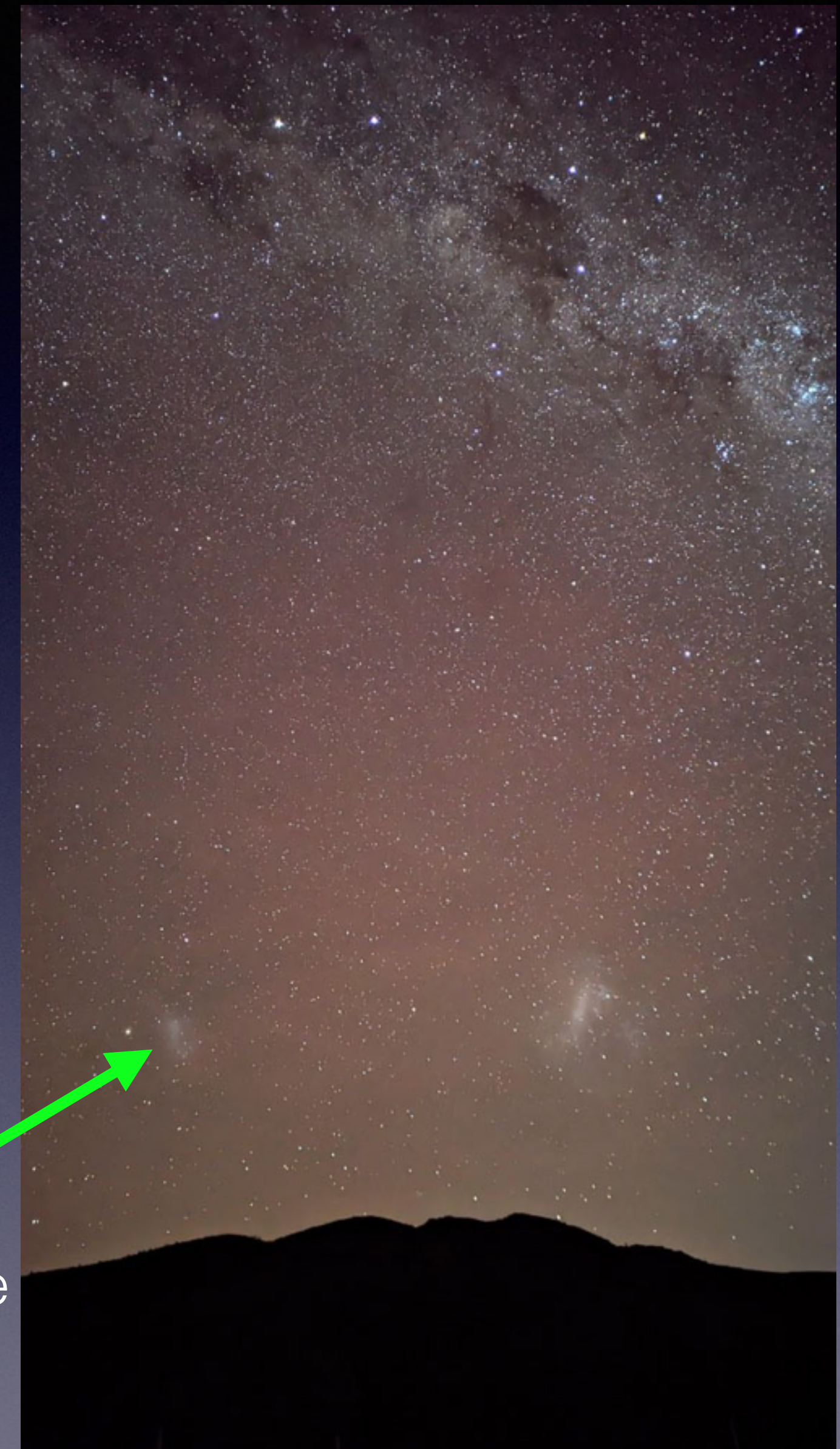


Exemple basse résolution sur une Be extragalactique

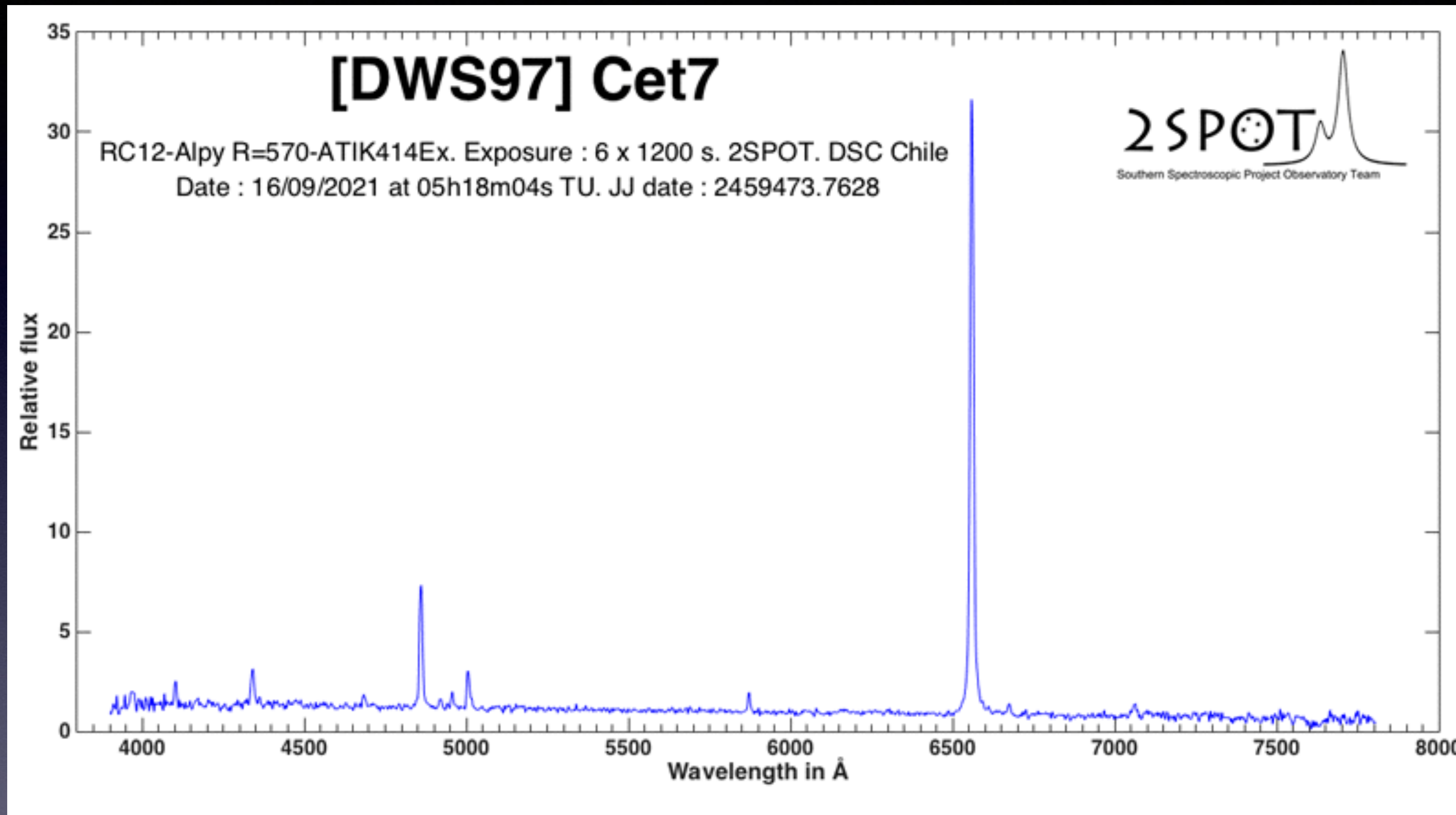


SMC 25 : Dans le petit nuage de Magellan (Mag. V=14,45)

On a également réalisé des spectres de candidates NP dans cette galaxie proche



Quelques résultats

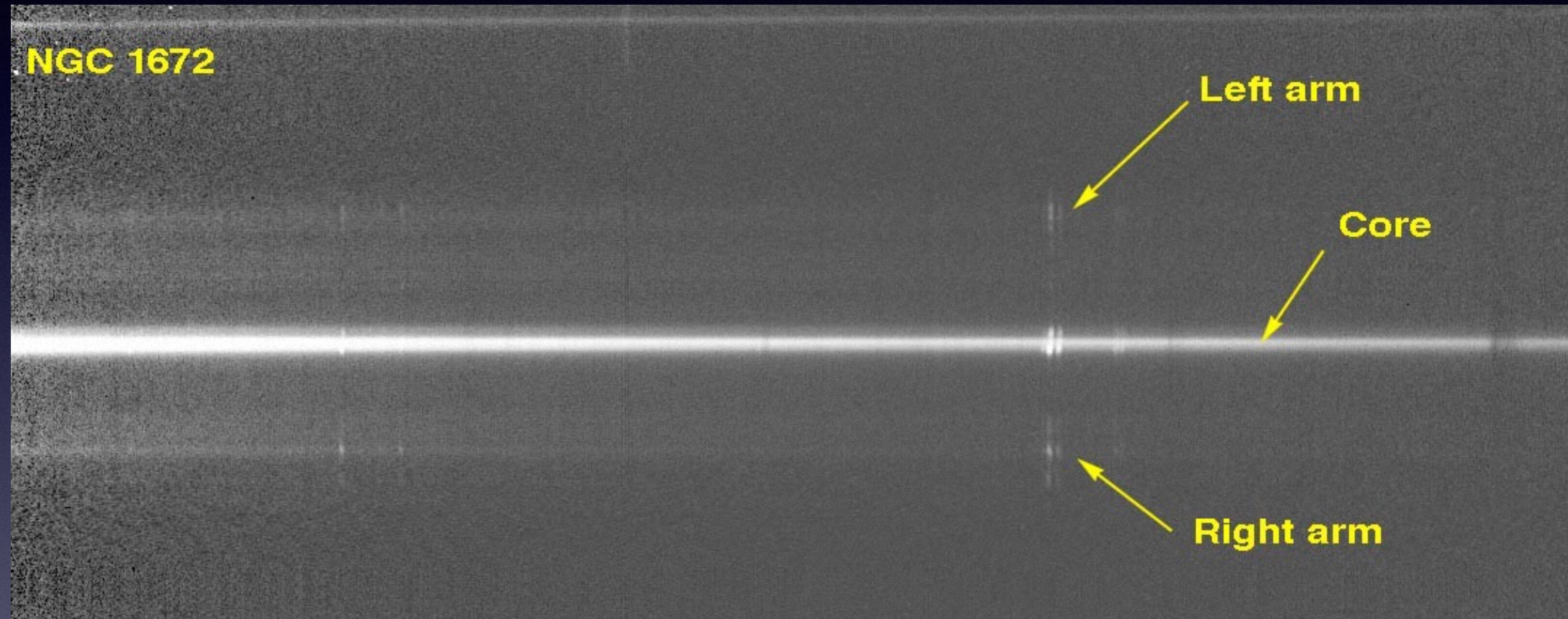


Etoile symbiotique **[DWS97] Cet7** (Magnitude V=15,14)

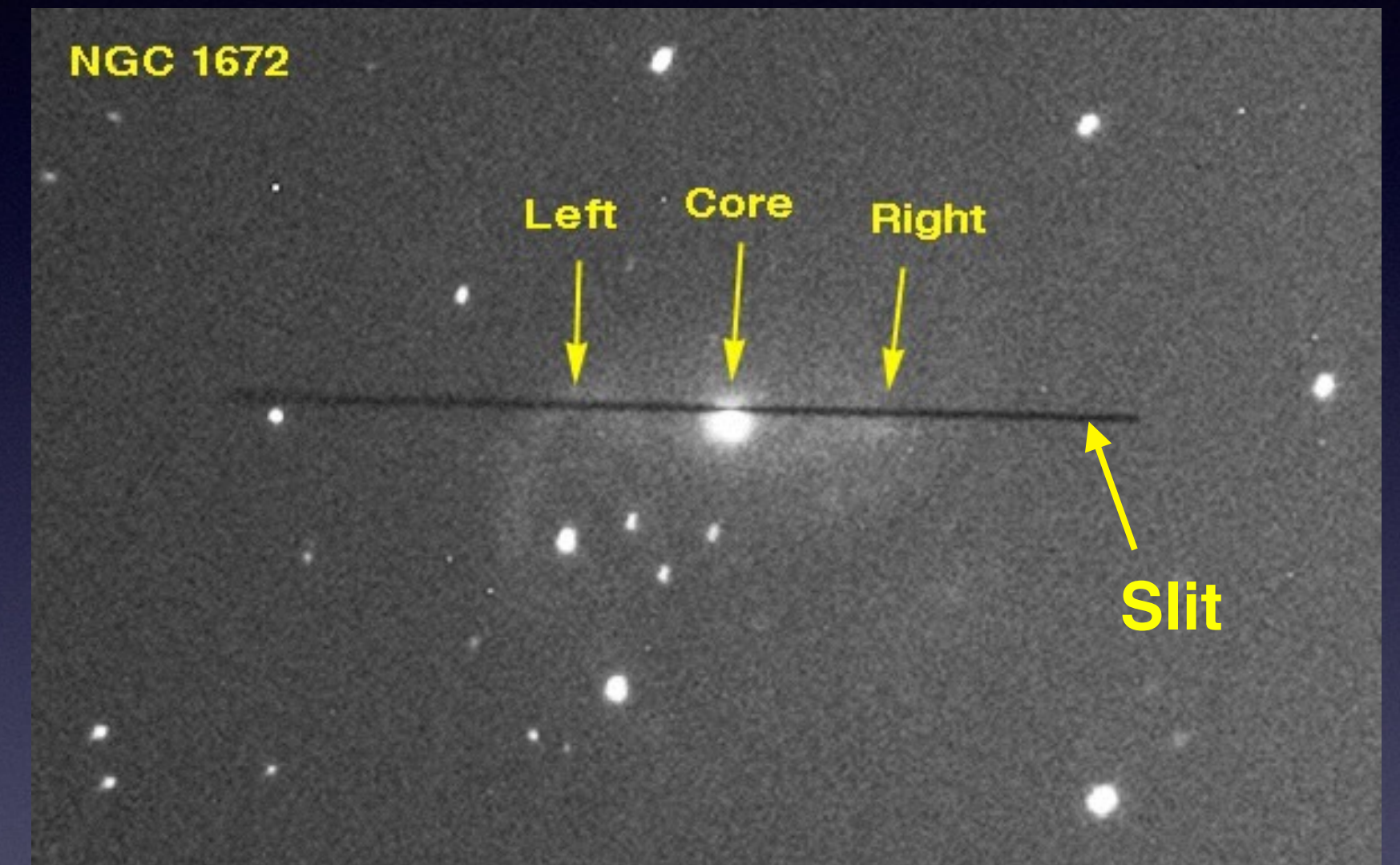


Quelques résultats

Galaxie de Seyfert : **NGC 1672**



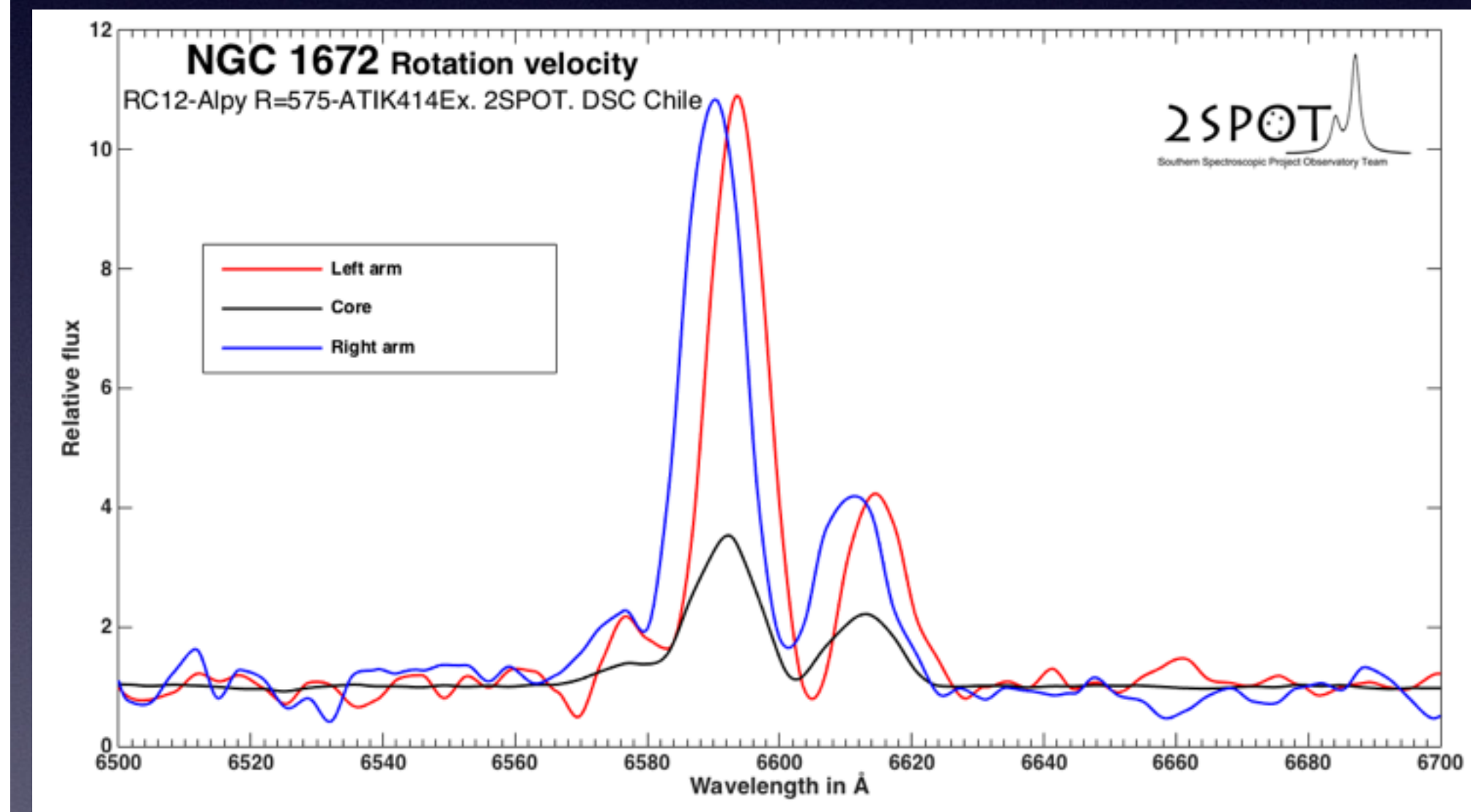
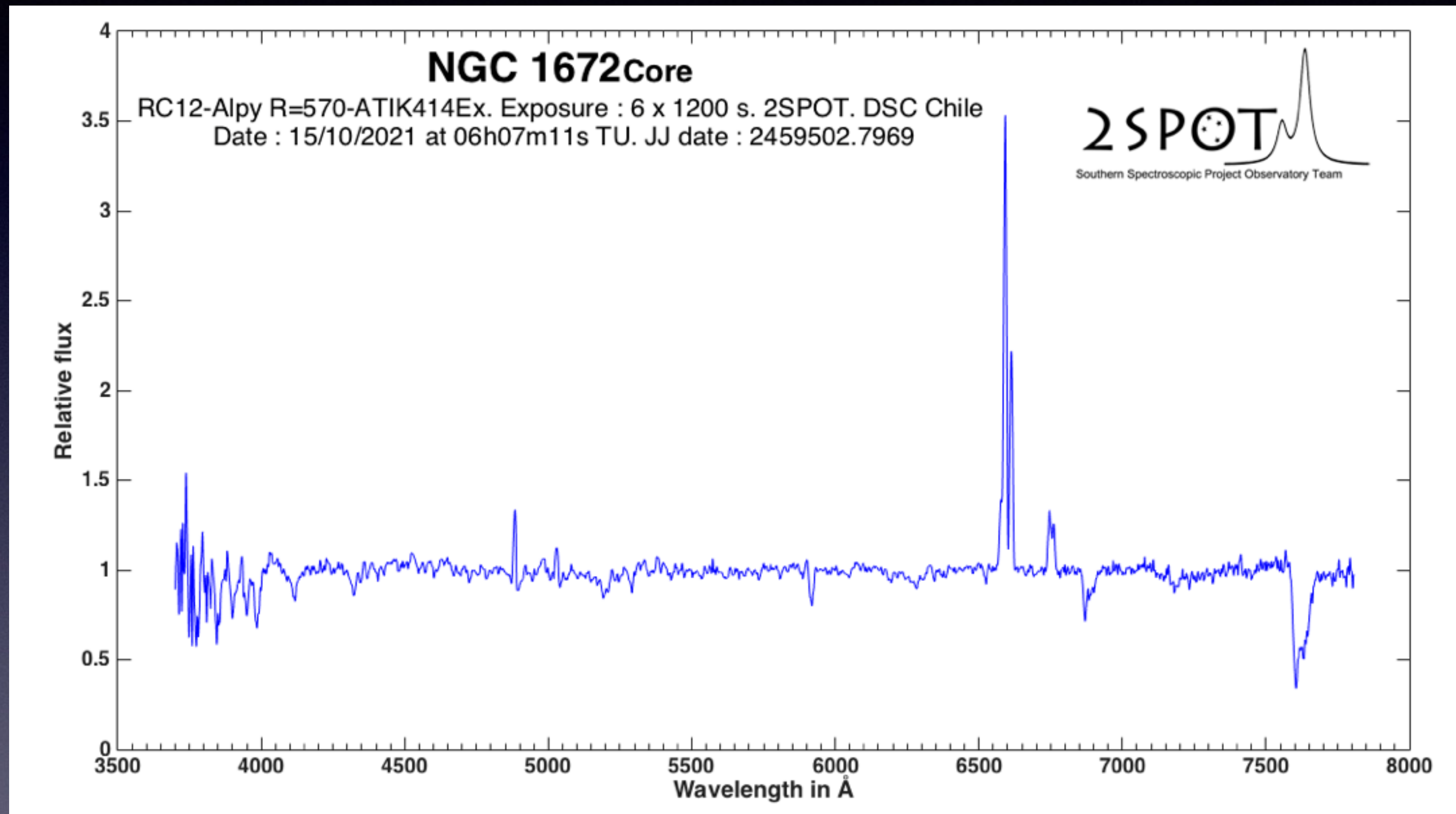
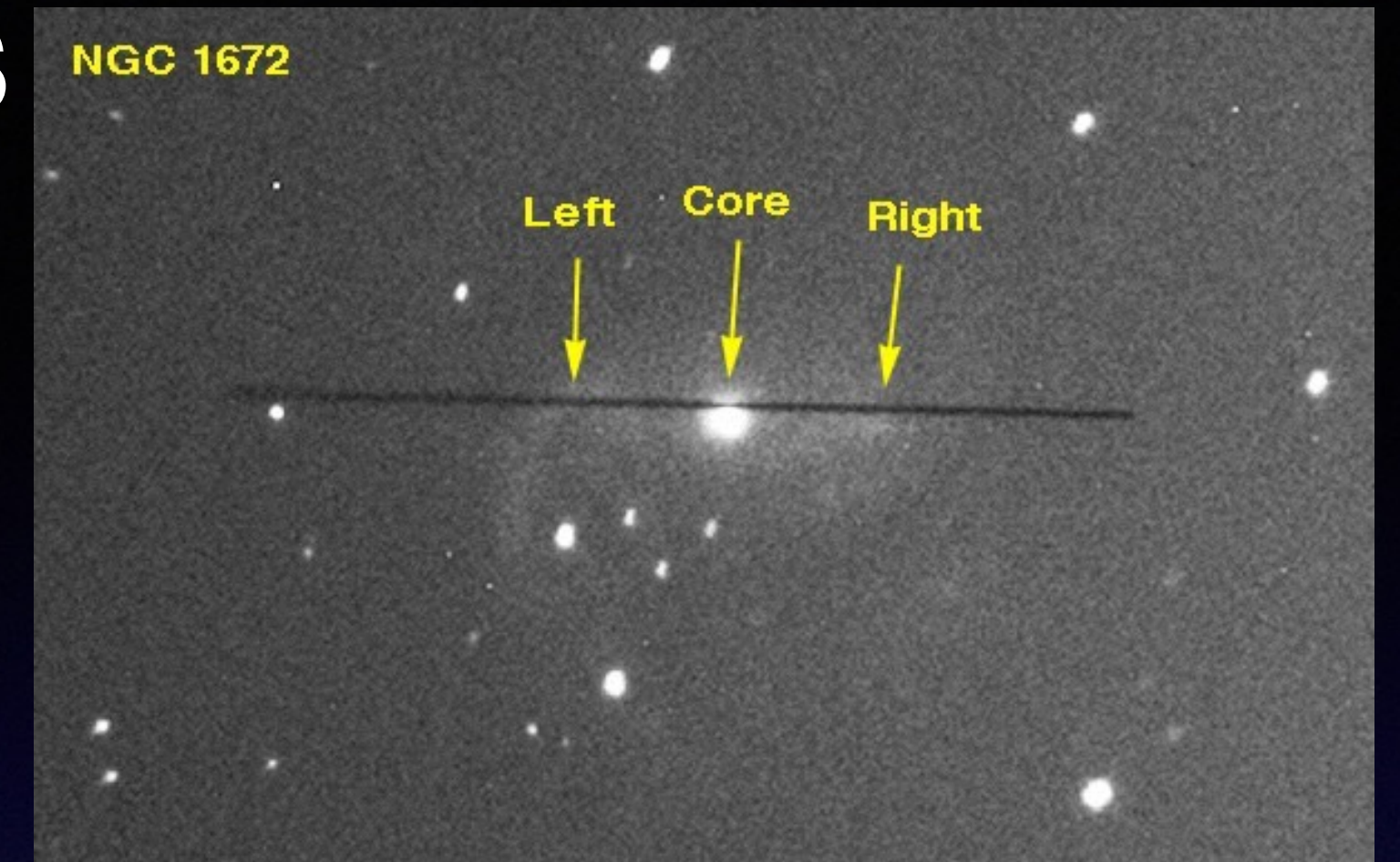
Spectre brut unitaire de 1200s



champ d'autoguidage
Position de la galaxie dans la fente

Quelques résultats

Galaxie de Seyfert : **NGC 1672**



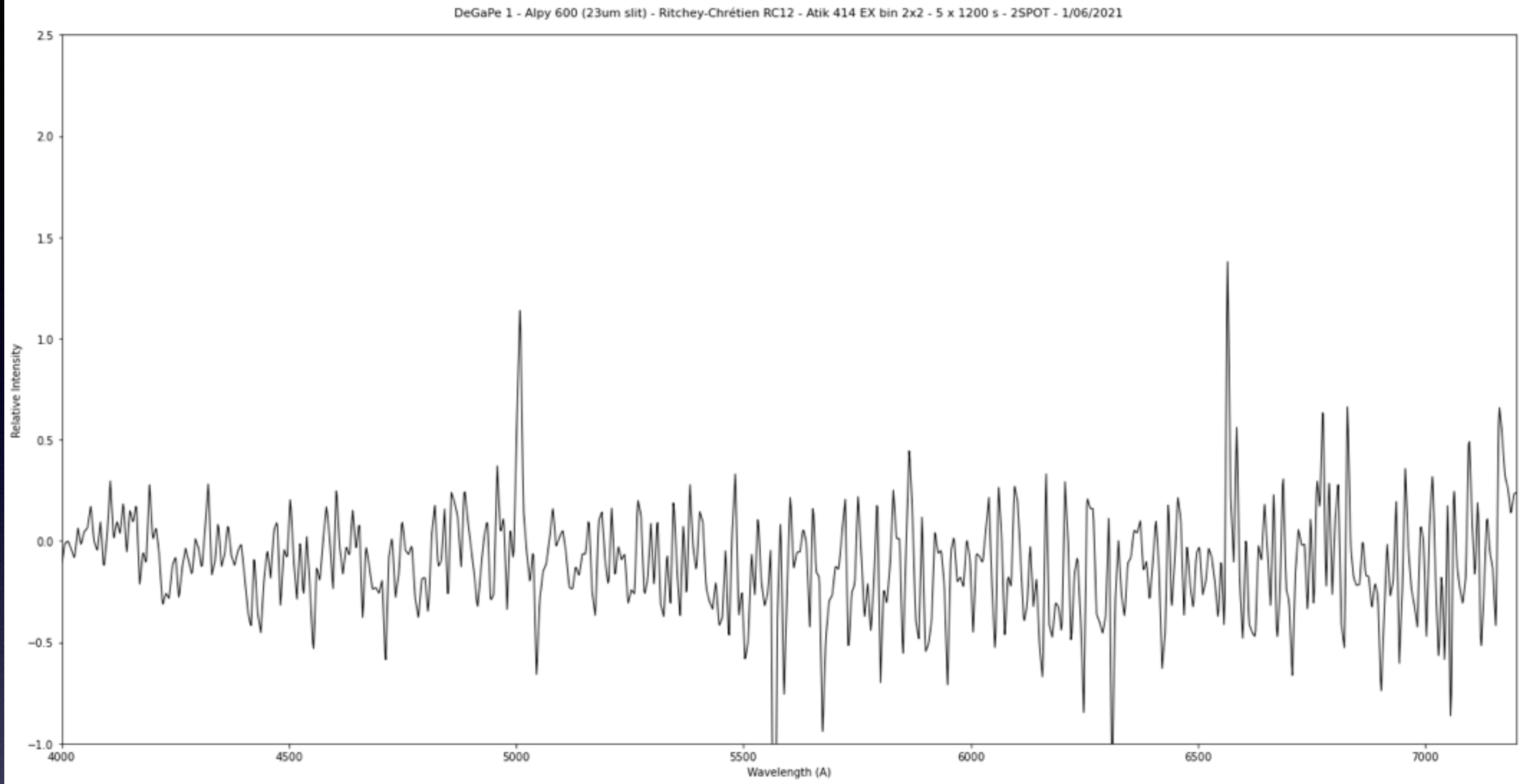
Vitesse radiale de 1335 km/s ($z=0,004464$)

Vitesse de rotation de la galaxie
Bras gauche en bleu
Bras droit en rouge

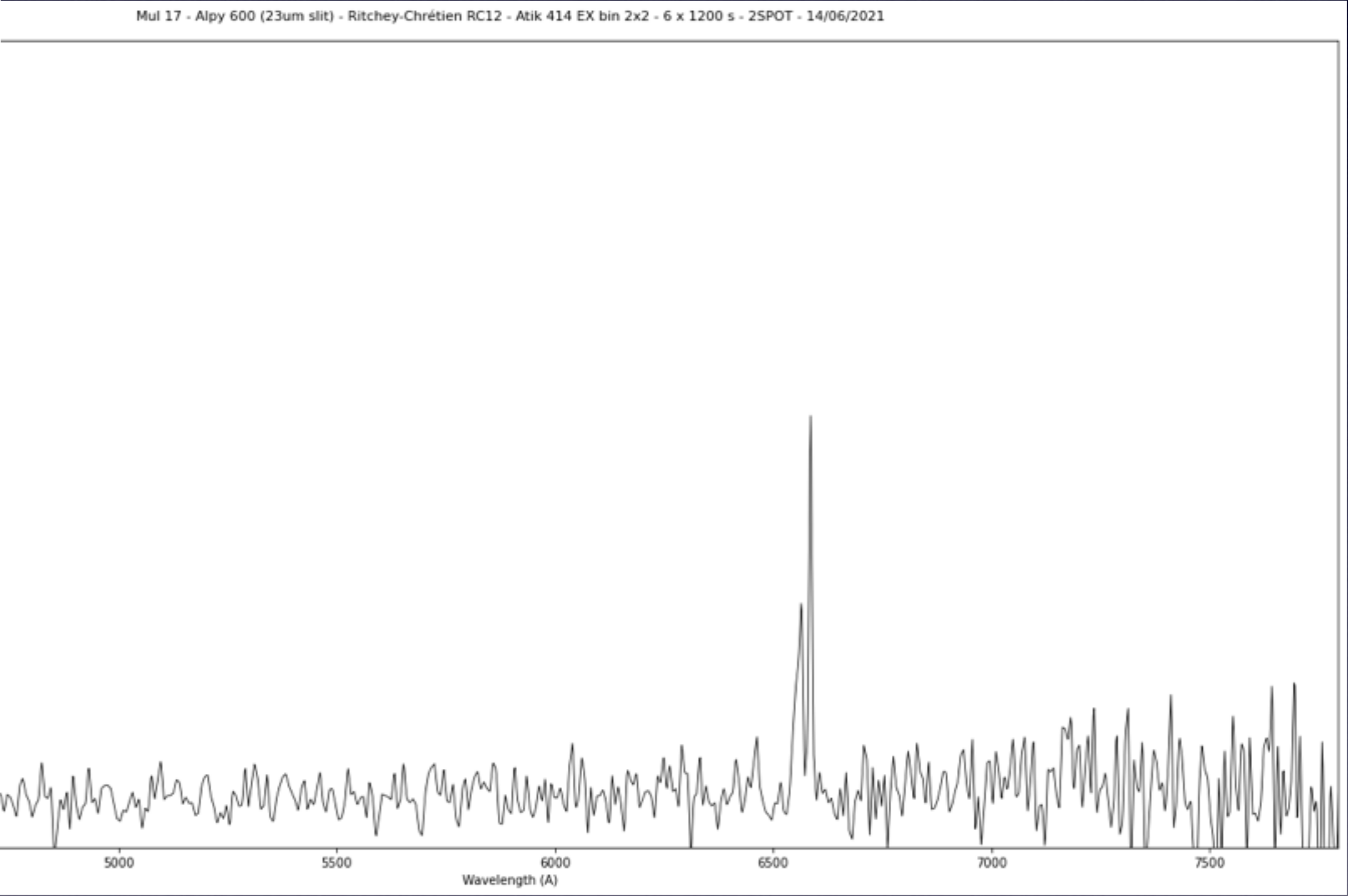


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Confirmation de candidates NP



DeGaPe 1



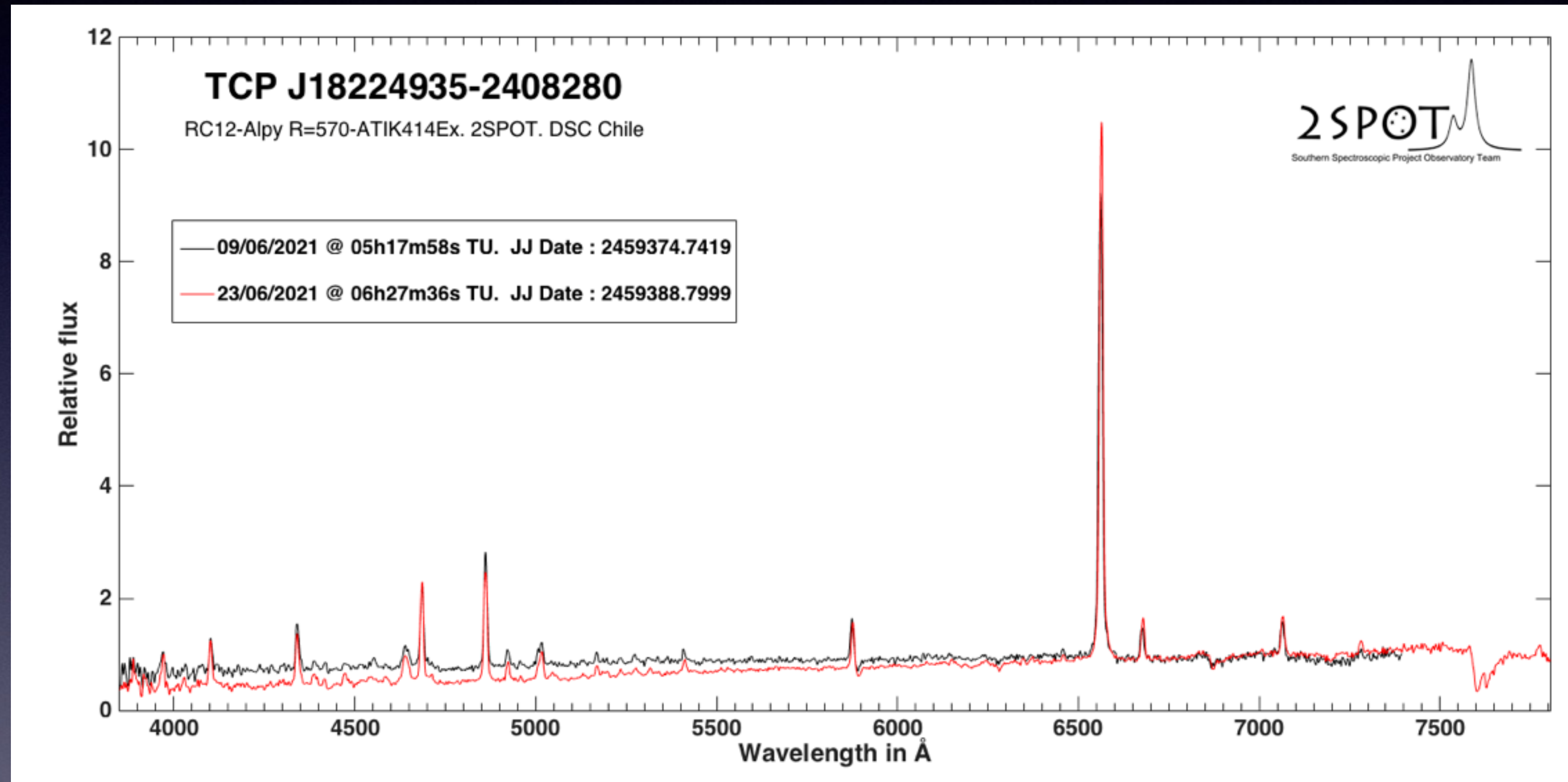
Mu1 17



Publications

(depuis le 24 Mai 2021)

UAI ATel #14691 (9 juin 2021)



TCP J18224935-2408280 is an outburst of a symbiotic star

ATel #14691; *J. Merc (UPJS in Kosice, Charles University), R. Galis (UPJS in Kosice), S. Charbonnel, O. Garde, P. L. Du, L. Mulato, T. Petit (2SPOT team)*
on 9 Jun 2021; 16:19 UT

Credential Certification: Jaroslav Merc (jaroslav.merc@student.upjs.sk)

Subjects: Infra-Red, Optical, Binary, Cataclysmic Variable, Transient, Variables

Referred to by ATel #: 14692, 14699

[Tweet](#)

The transient TCP J18224935-2408280 was discovered by Tadashi Kojima (CBAT "Transient Object Followup Reports"). Possible symbiotic nature was later suggested by Patrick Schmeer as he noticed that there is a Gaia DR2 LPV with a reported period of roughly 800 days 2" away from the announced position (Gaia DR2 4089297564356878720). The star is also included in the catalog of large-amplitude variables of Mowlavi et al. (2021, A&A 648, A44).

The light curve of TCP J18224935-2408280 from the ASAS-SN survey (Shappee et al., 2014, ApJ 788, 48; Kochanek et al., 2017, PASP, 129, 104502) shows 2.2 mag brightening which started between May 13, 2021 (JD 2 459 348.4) and May 16, 2021 (JD 2 459 351.4). The brightness reached the first maximum in about 5 days followed by a 10-day-long slight decrease of brightness by 0.5 mag. Since the beginning of June, the brightness is gradually rising again. No other brightenings are seen in the ASAS-SN light curve since March 14, 2016 (JD 2 457 461.8).

We obtained an optical spectrum of TCP J18224935-2408280 on June 9, 2021 (JD 2 459 374.7) using an Alpy600 spectrograph mounted at a remotely controlled 35-cm Ritchey-Chretien telescope located in Chile. The spectrum shows strong emission lines of H I, He I, [O III], and He II in addition to the K5-M0 continuum. The measured equivalent widths are 114, 30, 12, and 19 Ang for H alpha, H beta, H gamma, and He II 4686 emission lines, respectively.

The map of Schlafly & Finkbeiner (2011, ApJ 737, 103) gives total galactic extinction in the direction of TCP J18224935-2408280 $E(B-V) = 0.55$ corresponding to visual extinction of $A_V = 1.71$ mag. The parallax of the object in Gaia DR3 (Gaia Collaboration, 2021, A&A, 649, A1) is unreliable (0.0571 ± 0.0242 mas). Bailer-Jones et al. (2021, AJ, 161, 147) obtained a distance of around 8 kpc for the source. At such a distance, the apparent magnitude of TCP J18224935-2408280 is roughly consistent with class III stars.

Together with the ongoing outburst of the object and the long-term variability, these results very strongly indicate the symbiotic nature of TCP J18224935-2408280. The infrared colors of the object are consistent with an S-type symbiotic star.

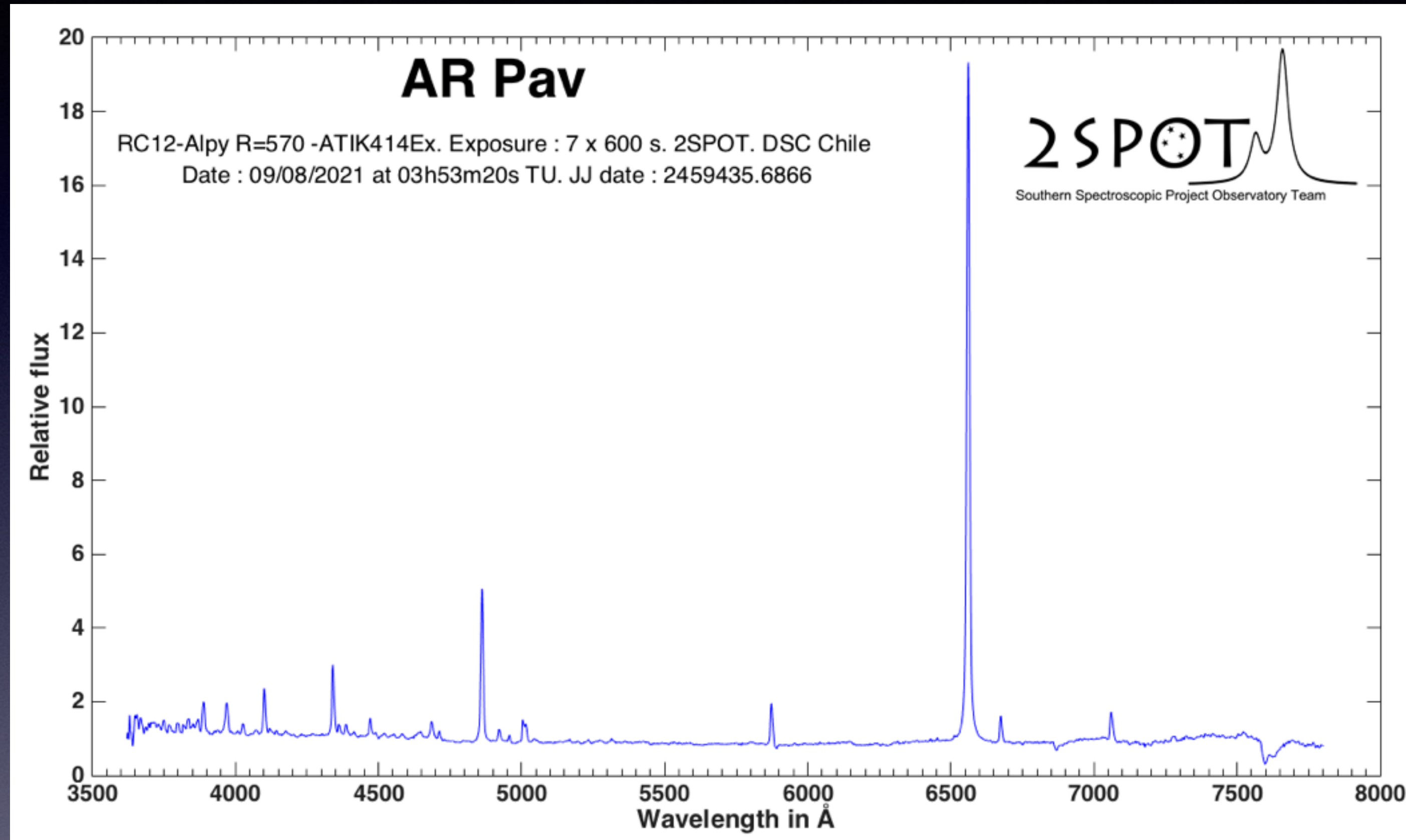
Southern Spectroscopic Project Observatory Team (2SPOT)



Rencontres du Ciel et de l'Espace 2024

Publications

(depuis le 24 Mai 2021)



UAI ATel #14837 (9 Août 2021)

Rencontres du Ciel et de l'Espace 2024

Ongoing brightening of the eclipsing symbiotic star AR Pav

ATel #14837; *J. Merc (UPJS in Kosice, Charles University), R. Galis (UPJS in Kosice), O. Garde, L. Mulato, P. Le Du, S. Charbonnel, T. Petit (2SPOT team)*
on 9 Aug 2021; 13:13 UT

Credential Certification: Jaroslav Merc (jaroslav.merc@student.upjs.sk)

Subjects: Infra-Red, Optical, Binary, Transient, Variables

[Tweet](#)

AR Pav is an eclipsing symbiotic star with an orbital period of around 605 days (e.g., Sekeráš et al., 2019, CAOSP, 49, 19, and references therein). Its photometric activity has been monitored since 1889 (see Fig. 1 in Skopal et al., 2001, IBVS, 5195 and Fig. 23 in Sekeráš et al., 2019, CAOSP, 49, 19). Its long-term photographic/B-band light curve is characterized by ~2 mag eclipses and an out-of-eclipse variability between 10 – 12 mag. Several brightenings of the system have been detected in past, with the most prominent outbursts observed in 1900 and 1935, when the star reached ~9 mag. In the recent decade, the star experienced only smaller brightenings in 2013 and 2014 during which the star was always fainter than 10.5 mag.

The **photometric observations of AR Pav by the ASAS-SN survey** (Shappee et al., 2014, ApJ 788, 48; Kochanek et al., 2017, PASP, 129, 104502) revealed an ongoing brightening by about 1.2 mag (with a peak brightness around 10 mag in the g filter). A slight rise towards the maximum started in April 2021 and accelerated in May 2021, reaching the maximal brightness at the beginning of June. Since then, according to the **recent ASAS-SN light curve** of AR Pav, brightness seems to fluctuate in a pattern similar to superhumps observed in cataclysmic variables (although some additional unexplained scatter seems to be present in the light curve). Such behavior is not detected during the rise to the maximum nor during the previous orbital cycles observed by ASAS-SN.

We obtained an **optical spectrum of AR Pav** on August 9, 2021 (JD 2 459 435.7) using an Alpy600 spectrograph mounted at a remotely controlled 35-cm Ritchey-Chretien telescope located in Chile. The spectrum shows a weak continuum of an M giant in addition to strong emission lines of H I, He I, [O III], relatively faint He II, and several emission lines of Fe II. Emission lines with a higher ionization potential (e.g., [Fe VII]) are not detected in our spectrum. The overall appearance of the spectrum is consistent with a Z And-type symbiotic star in an outburst.

Further multi-band photometric and spectroscopic observations are encouraged.

Southern Spectroscopic Project Observatory Team (2SPOT)



Publications

(depuis le 24 Mai 2021)

Brightening of SS73 141: Outburst of Z And-type?

ATel #14874; *J. Merc (UPJS in Kosice, Charles University), R. Galis (UPJS in Kosice), P. Velez (ARAS Group), S. Charbonnel, O. Garde, P. Le Du, L. Mulato, T. Petit (2SPOT team, ARAS Group)*

on 26 Aug 2021; 17:58 UT

Credential Certification: Jaroslav Merc (jaroslav.merc@student.upjs.sk)

Subjects: Infra-Red, Optical, Binary, Transient, Variables

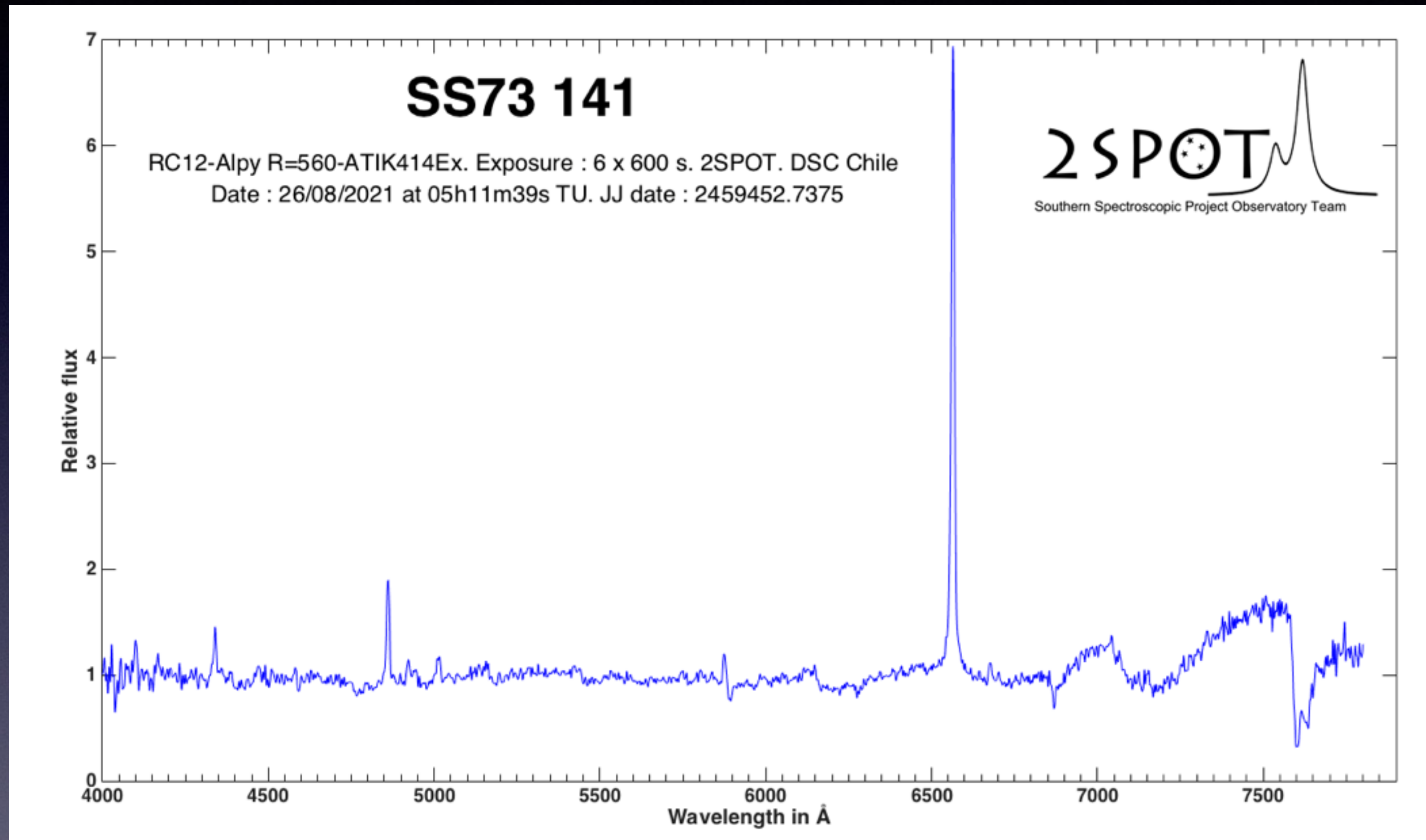
 Tweet

SS73 141 (=WRAY 16-384) is a poorly studied symbiotic star. The optical spectrum of the object, obtained on June 17, 1978, and presented by Allen (1984, PASA, 5, 369) showed prominent TiO bands of late-type giant (later classified as M5 by Murset & Schmid, 1999, A&AS, 137, 473) together with strong emission lines of H I and He II, and relatively faint emission lines of neutral helium.

Recently, Tadashi Kojima (Gunma, Japan) reported a brightening of SS73 141 (see [vsnet-alert 26105](#)). In [vsnet-alert 26109](#), Taichi Kato hypothesized, that the ongoing brightening of SS73 141 might be an outburst of a slow symbiotic nova, as opposed to an outburst of classical symbiotic star (Z And-type). For a description of individual types of outbursts observed in symbiotic stars, see, e.g., recent review of symbiotic stars by Munari (2019, arXiv:1909.01389).

According to recent [ASAS-SN light curve](#) (Shappee et al., 2014, ApJ 788, 48; Kochanek et al., 2017, PASP, 129, 104502), the outburst started at the beginning of May 2021 and the brightness reached the maximum in the middle of July (around $g = 13.4$). The median magnitude of the object before the outburst was $g \sim 15$ resulting in the amplitude of the outburst of 1.6 mag (in g). Since the middle of July, the brightness has remained more or less constant. No previous brightenings are seen in the [ASAS-SN light curve](#) nor are reported in the literature (up to our knowledge), making the ongoing brightening the first recorded outburst of SS73 141.

We obtained [two low-resolution optical spectra of the object](#) at JD 2 459 451.47 (August 25, 2021) and JD 2 459 452.24 (August 26, 2021). They cover the wavelength range of 4000-5639 Å and 3900-7802 Å, respectively. The spectra revealed the Balmer lines in emission, together with several faint emission lines of neutral helium. Emission lines with higher ionization potential (e.g., He II) are not detected in our spectra. TiO bands are rather weak, especially when compared with the spectrum presented by Allen (1984).



UAI ATel #14874 (26 Août 2021)

Rencontres du Ciel et de l'Espace 2024



Publications



(depuis le 24 Mai 2021)

RNAAS (Mars 2022)

RNAAS RESEARCH NOTES OF THE AAS

OPEN ACCESS

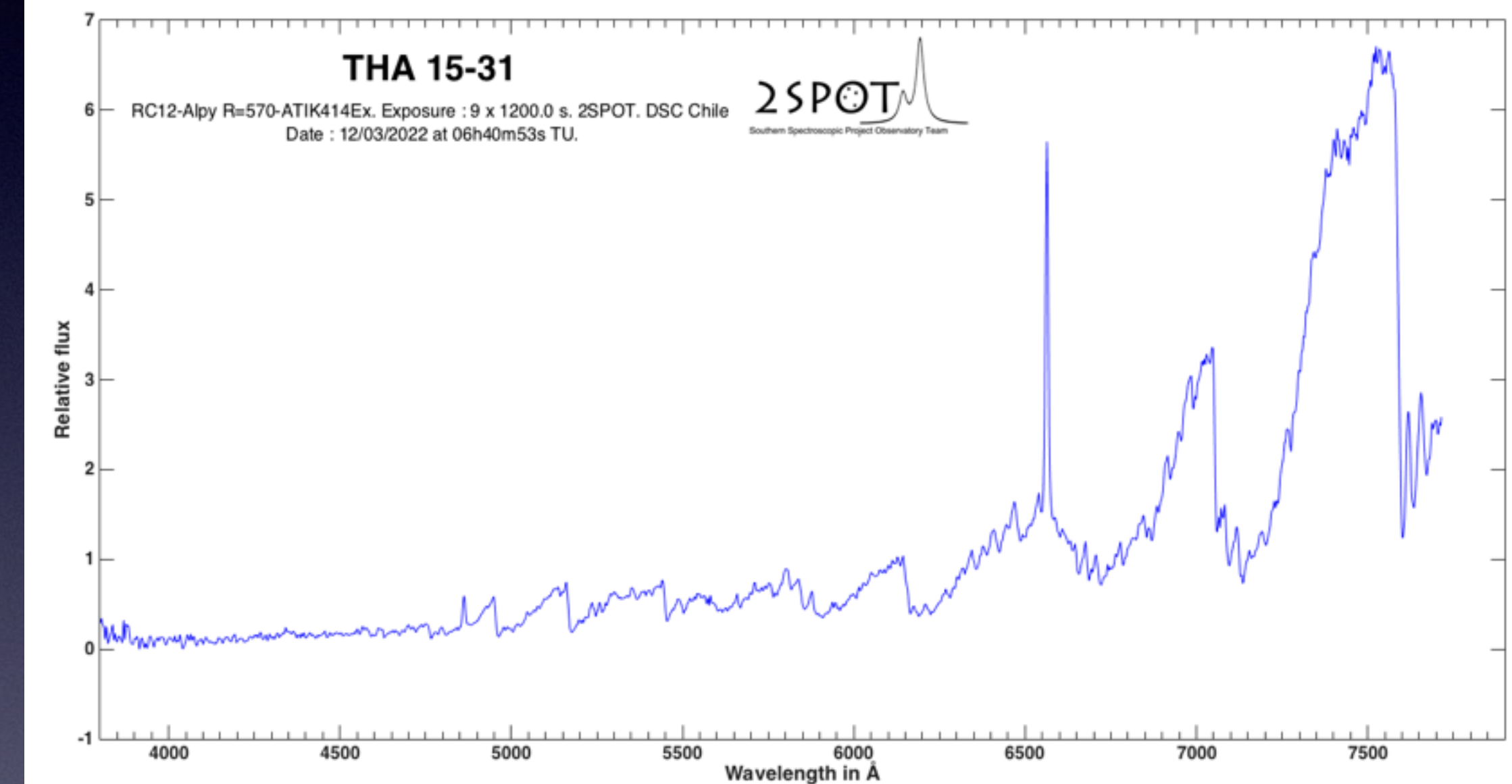
Low-resolution Optical Spectroscopy of Recently Discovered Accreting-only Symbiotic Star THA 15–31

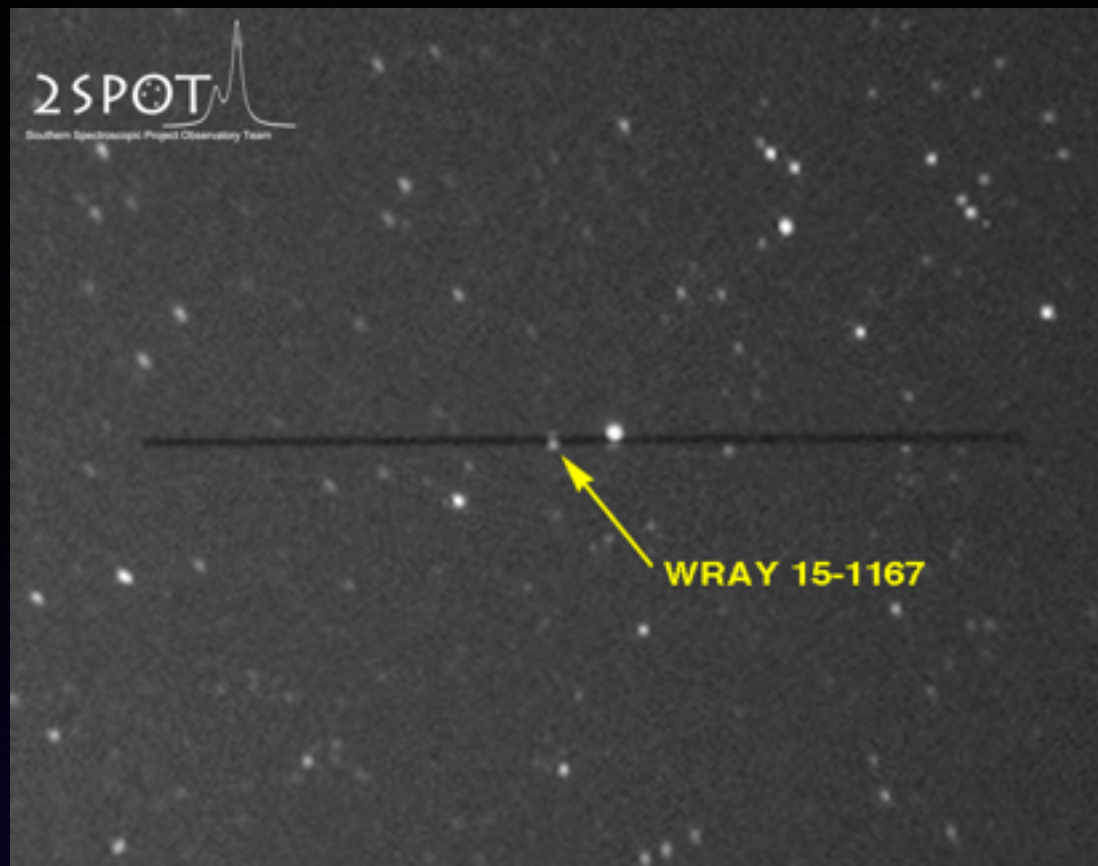
Jaroslav Merc^{1,2} , Stéphane Charbonnel^{3,4}, Olivier Garde^{3,4}, Pascal Le Du^{3,4}, Lionel Mulato^{3,4}, Thomas Petit^{3,4}, and Rudolf Gális² 

Published March 2022 • © 2022. The Author(s). Published by the American Astronomical Society.

[Research Notes of the AAS, Volume 6, Number 3](#)

Citation Jaroslav Merc et al 2022 Res. Notes AAS 6 54





Publications

(depuis le 24 Mai 2021)

Gaia22bou: First recorded outburst of symbiotic star WRAY 15-1167

ATel #15340; *J. Merc (UPJS In Kosice, Charles University), P. Velez, H. Barker (ARAS Group), S. Charbonnel, O. Garde, P. Le Du, L. Mulato, T. Petit (2SPOT team, ARAS Group), R. Galls (UPJS In Kosice)*

on 20 Apr 2022; 12:56 UT

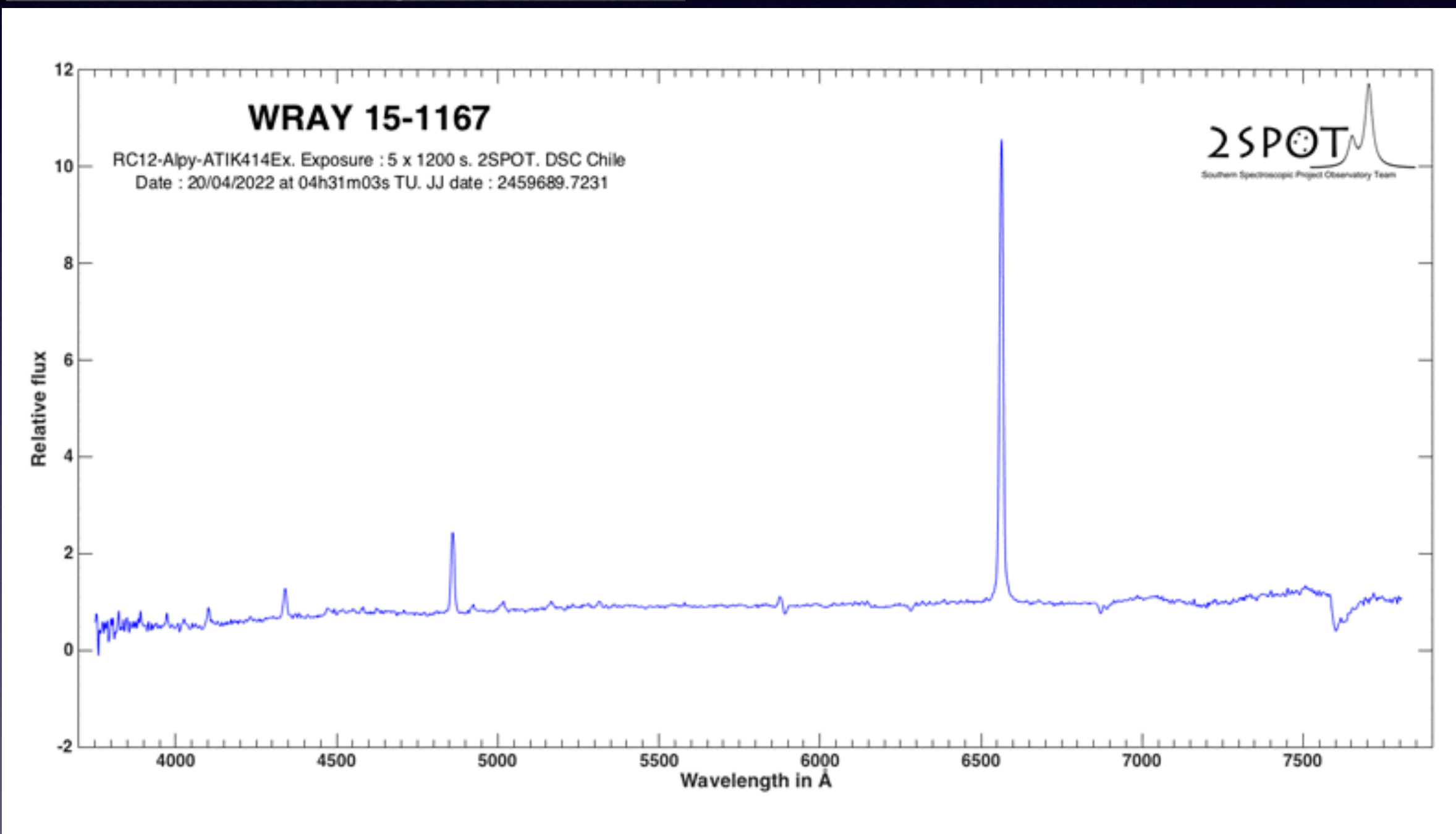
Credential Certification: Jaroslav Merc (jaroslav.merc@student.upjs.sk)

Subjects: Optical, Binary, Transient, Variables

Tweet

WRAY 15-1167 is a symbiotic star discovered by Miszalski & Mikolajewska (2014, MNRAS, 440, 1410) in a sample of emission-line objects selected from the AAO/UKST SuperCOSMOS H α Survey. This target was classified as a symbiotic star based on the spectrum obtained on June 18, 2013, with the Southern African Large Telescope (SALT) that showed the continuum of M4 giant together with strong Balmer lines, He I, He II, and possibly also [Fe VII] in emission.

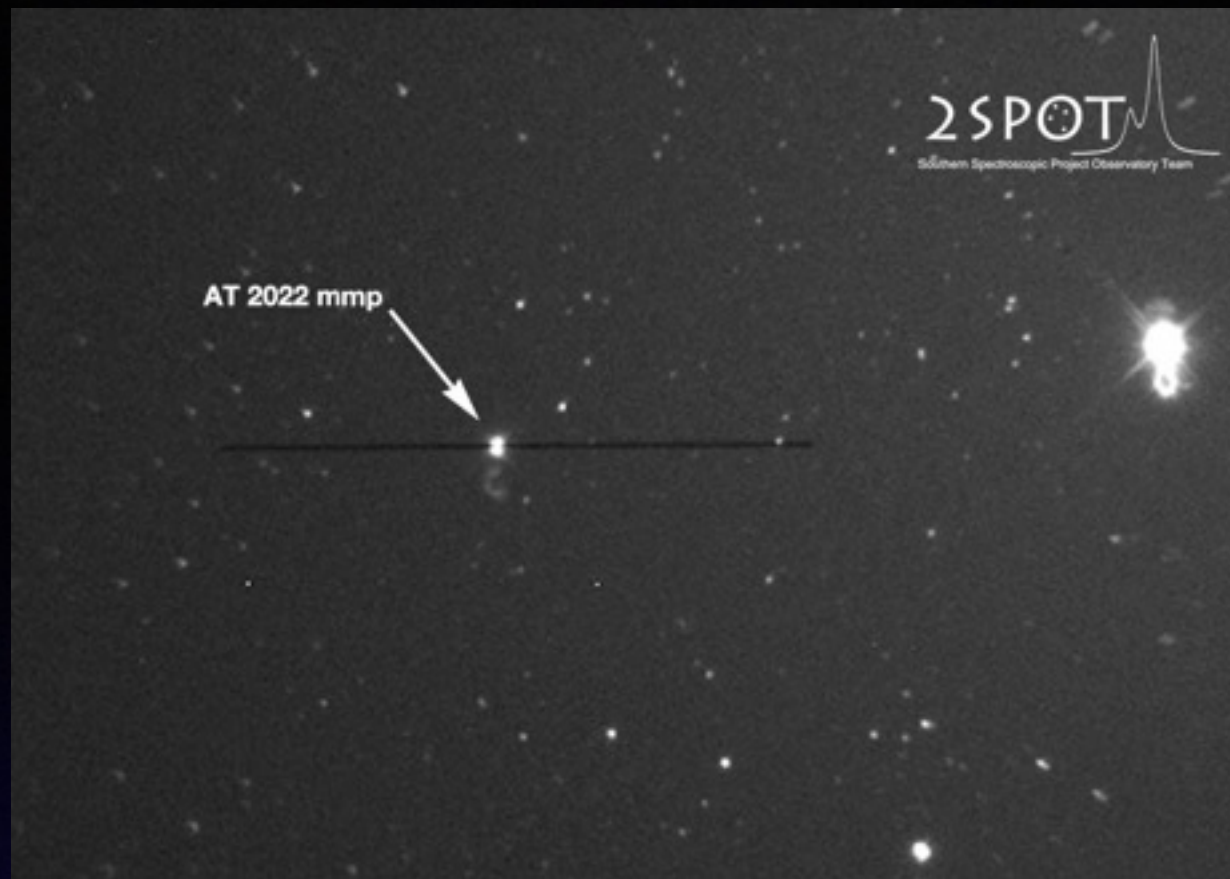
The brightening of WRAY 15-1167 was detected by the Gaia satellite and announced as a Gaia Science Alert (Hodgkin et al., 2021, A&A, 652, A76) on April 16, 2022 (**Gaia22bou**), when the star had the magnitude $G = 12.66$. The Gaia G light curve showed that the outburst probably started much earlier (at the turn of years 2021 and 2022). The measurement from the end of February revealed the system to be even brighter ($G = 12.58$) than at the time of alert publication. The pre-outburst magnitude of WRAY 15-1167 was around $G = 14$. This suggests the outburst amplitude of around 1.4 mag in the Gaia G filter.



UAI ATel #15340 (20 avril 2022)



Rencontres du Ciel et de l'Espace 2024



Publications

(depuis le 24 Mai 2021)

Spectroscopic monitoring of bright Galactic nova ASASSN-22hw

ATel #15435; *J. Merc (UPJS In Kosice, Charles University), H. Barker, P. Velez (ARAS Group), S. Charbonnel, O. Garde, P. Le Du, L. Mulato, T. Petit (2SPOT Team, ARAS Group), T. Love (ARAS Group), R. Galls (UPJS In Kosice)*

on 15 Jun 2022; 09:12 UT

Credential Certification: Jaroslav Merc (jaroslav.merc@student.upjs.sk)

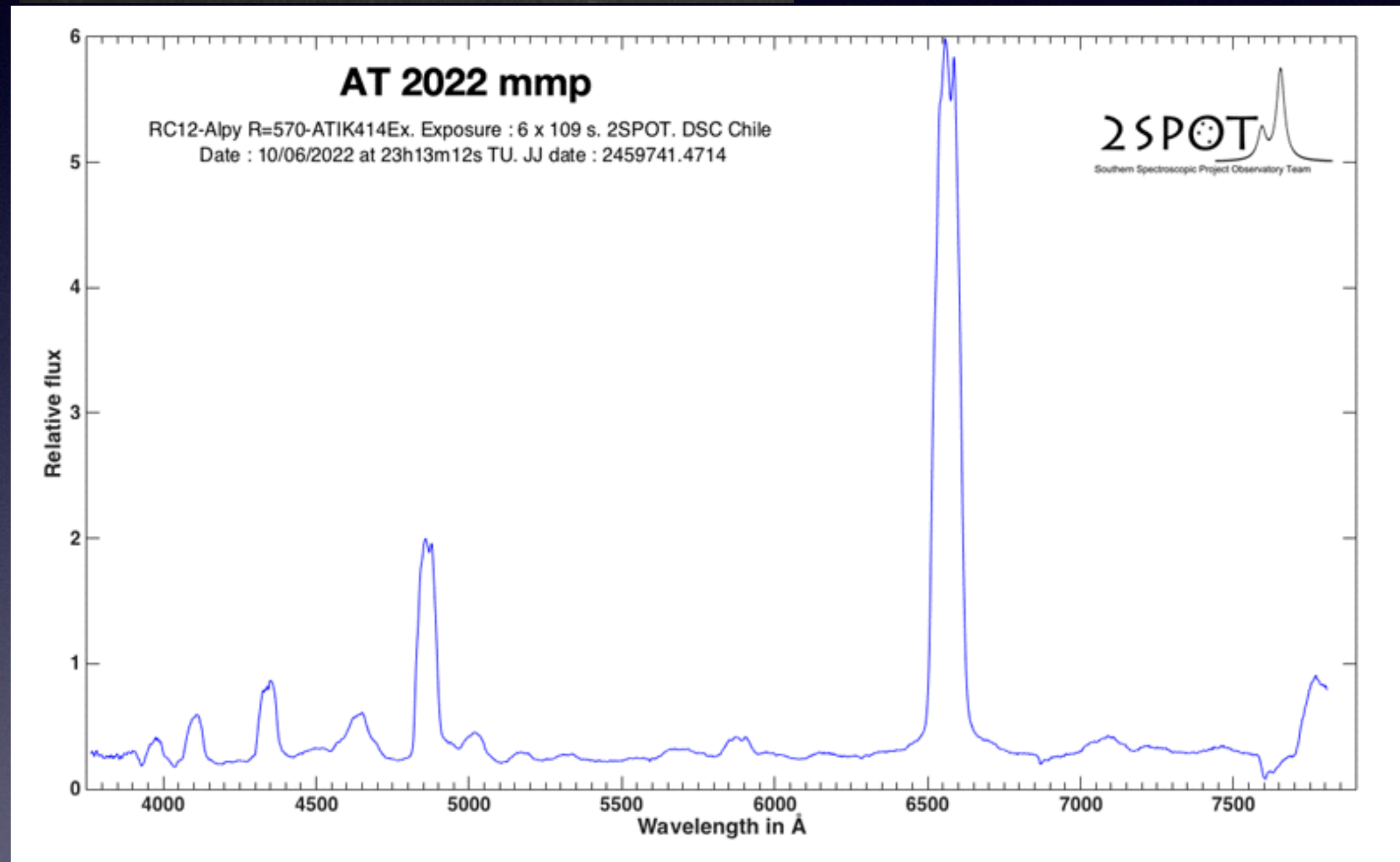
Subjects: Optical, Nova, Transient



The discovery of ASASSN-22hw (now also **AT 2022mmp** and Nova Mus 2022) was reported to the **Transient Name Server** by K. Z. Stanek on the behalf of the ASAS-SN team on June 10, 2022. The object was proposed to be a 'likely galactic nova' already in the decline. The first spectroscopic observations of one of us (H.B.), that were already reported to TNS on June 11, confirmed the nova nature of the transient. The request for subsequent observations was made through the **AAVSO Alert Notice 781**.

The recent **ASAS-SN light curve** of the source (Shappee et al., 2014, ApJ 788, 48; Kochanek et al., 2017, PASP, 129, 104502) shows the maximum brightness of $g = 8.7$ mag that occurred on June 8, 2022. In the five days after the first detection, the brightness declined by about 1.7 mag already. The minimum pre-outburst magnitude derived from the Gaia EDR3 data by Patrick Schmeer and reported to the International Variable Star Index was about $V = 19.8$ mag. The photometric follow-up is also available in the **AAVSO database**.

UAI ATel #15435 (15 juin 2021)



Publication sur la découverte d'une nouvelle étoile symbiotique

Cornell University

We gratefully acknowledge support from the Simons Foundation and member institutions.

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[Submitted on 29 Sep 2022]

DeGaPe 35: Amateur discovery of a new southern symbiotic star

Thomas Petit, Jaroslav Merc, Rudolf Gális, Stéphane Charbonnel, Thierry Demange, Richard Galli, Olivier Garde, Pascal Le Dû, Lionel Mulato

In this work, we present the discovery and characterization of a new southern S-type symbiotic star, DeGaPe 35. We have obtained the low-resolution spectroscopic observations and supplemented them with photometry from Gaia DR3 and other surveys. The optical spectra of this target show prominent emission lines, including highly ionized [Fe VII] and O VI lines. The cool component of this symbiotic binary is an M4-5 giant with effective temperature $\sim 3\,380 - 3\,470$ K and luminosity $\sim 3\,000 L_{\odot}$ (for the adopted distance of 3 kpc). The hot component is a shell-burning white dwarf. The photometric observations of the Gaia satellite, published recently in the Gaia DR3 suggested the variability with the period of about 700 - 800 days that we tentatively attributed to the orbital motion of the binary.

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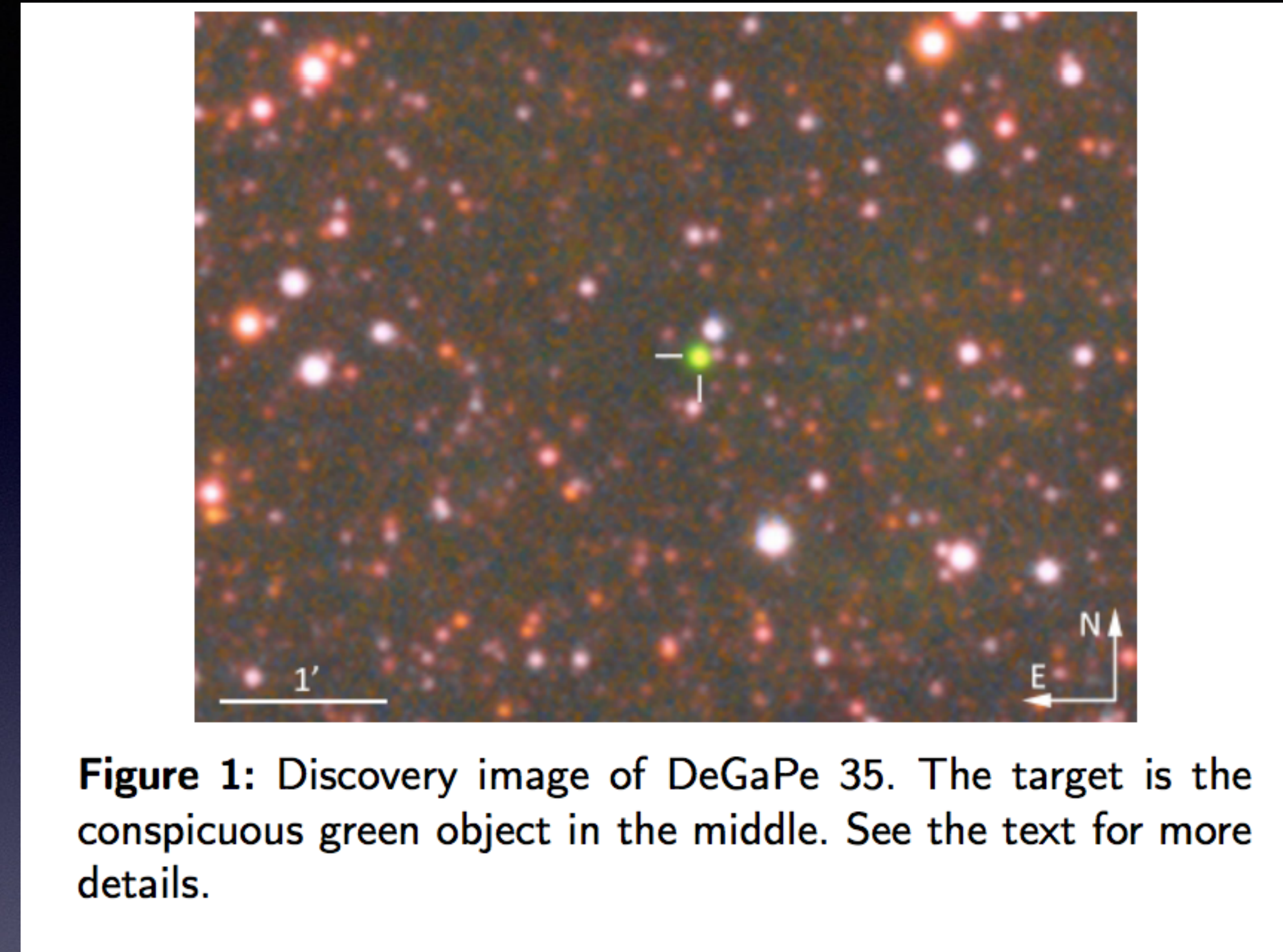
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2. Observational data

We have obtained two low-resolution spectra of DeGaPe 35 using the Southern Spectroscopic Project Observatory³ remotely-operated Ritchey-Chrétien 12" telescope located at the Deep Sky Chile facilities, equipped with an Alpy600 spectrograph (23 μm slit, providing a resolving power ~ 550) and Atik 414EX cooled camera. The first spectroscopic data were obtained on June 11, 2021 (JD

Une collaboration imagerie et spectro



Participation à l'obtention du **Prix Gemini 2022** de la SAF-SF2A pour nos travaux dans le cadre d'une collaboration pro/am de recherche et de confirmation de NP



Notre 2ème setup

Novembre 2022



Un 2ème spectrographe : **eShel**

- Spectrographe Echelle
- Fibre optique
- Couverture spectrale : 3900 Å to 7300 Å
- Resolution $R=11000$
- ATIK 460ex CCD camera
- Financé en partie par le CNRS/LESIA

Notre 2ème setup

Novembre 2022



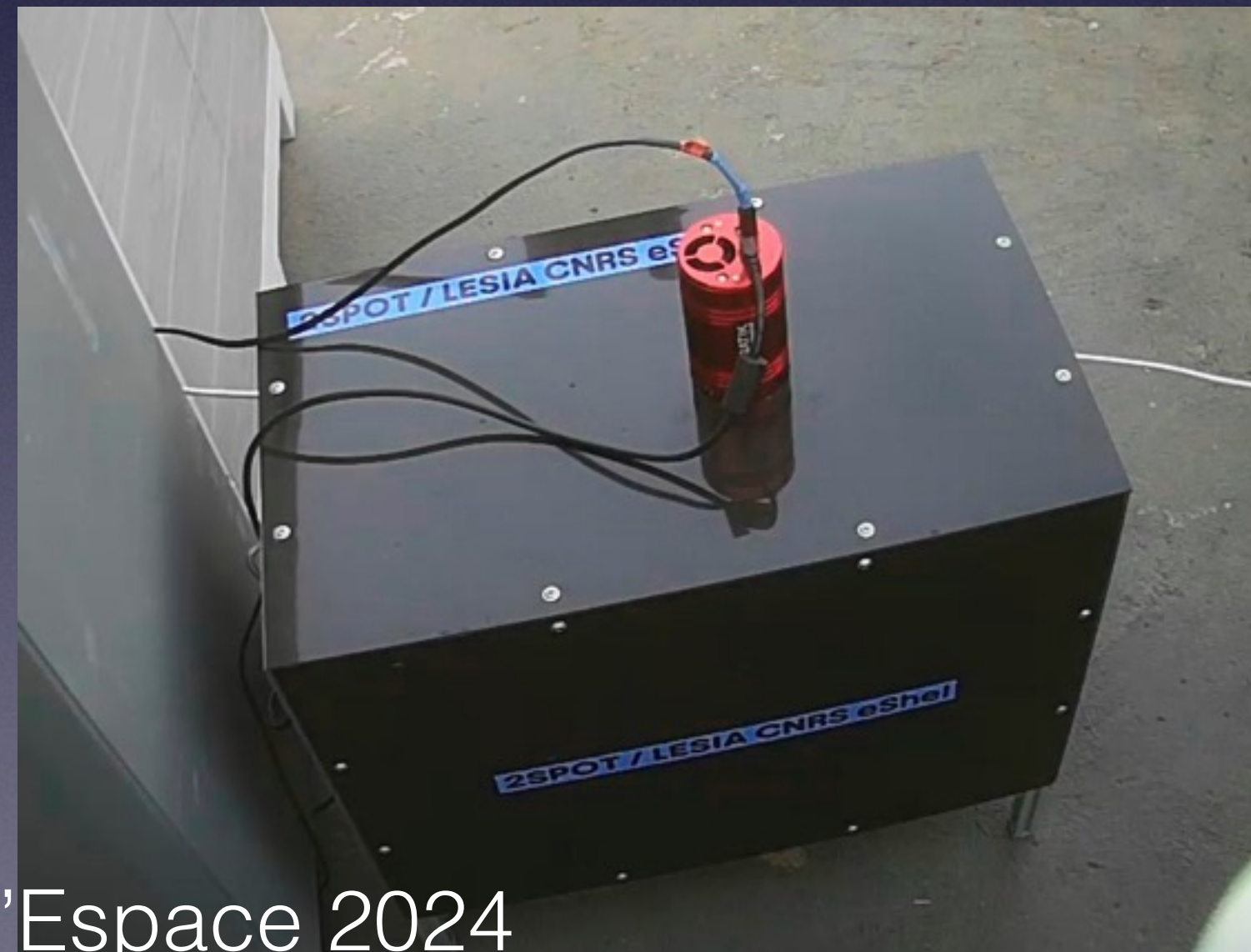
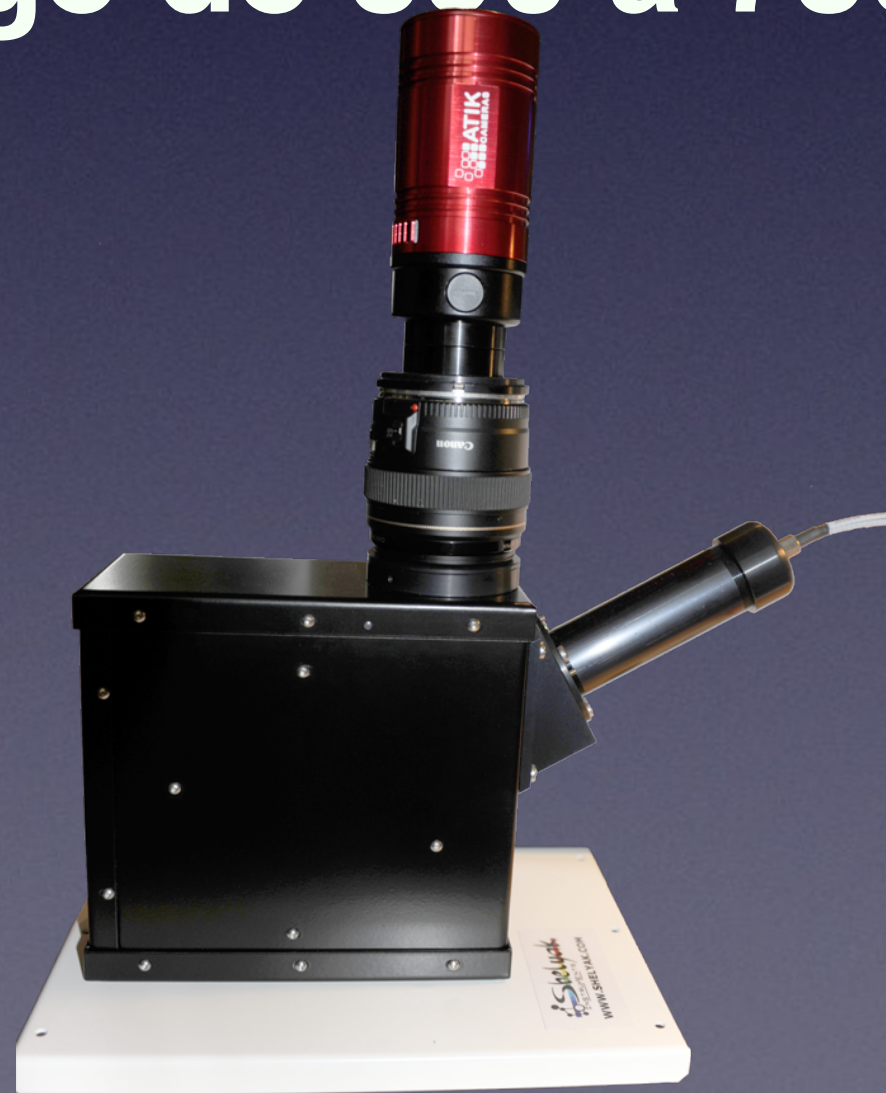
Départ France : 5 Août 2022
Arrivé Chili : 7 novembre 2022



Rencontres du Ciel et de l'Espace 2024

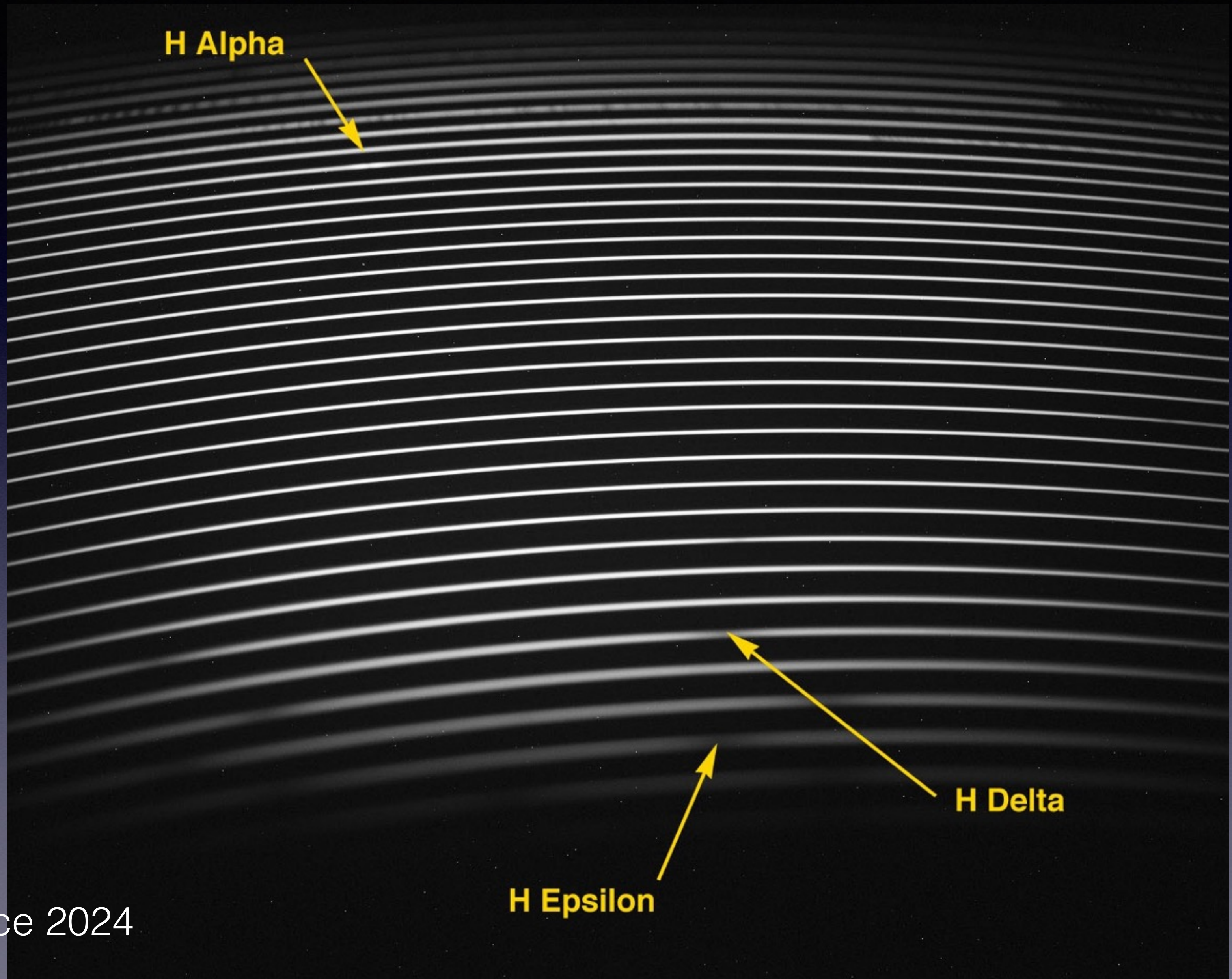
Nos **2** setups Spectro au Chile

- Télescope Newton de 305mm pour la basse résolution low-resolution avec un **Alpy 600**
Plage de 370 à 750 nm - R=600
- Télescope RC 305mm Ritchey-Chrétien Pour la moyenne résolution avec un **eShel**
Plage de 390 à 730 nm - R=11000



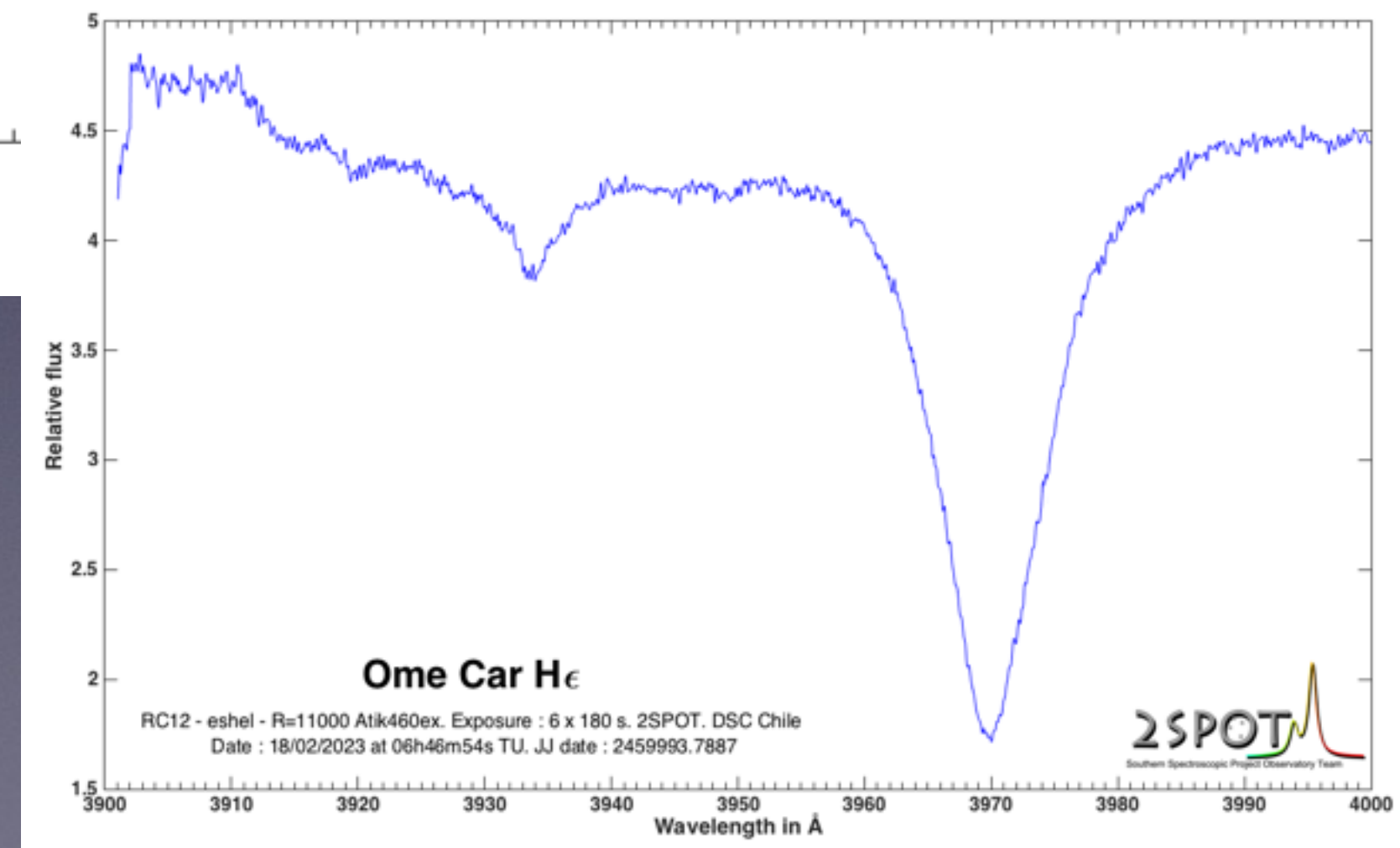
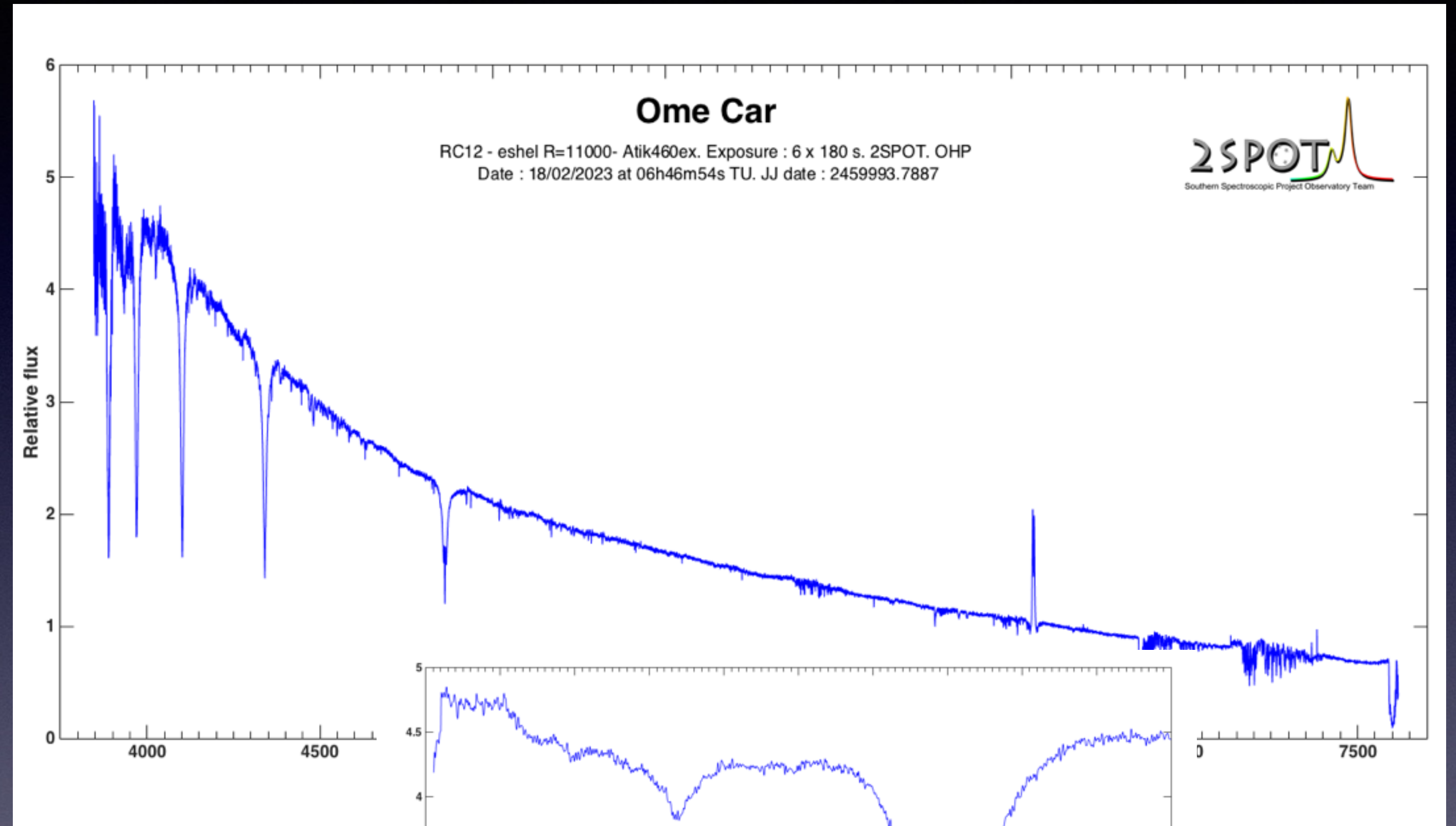
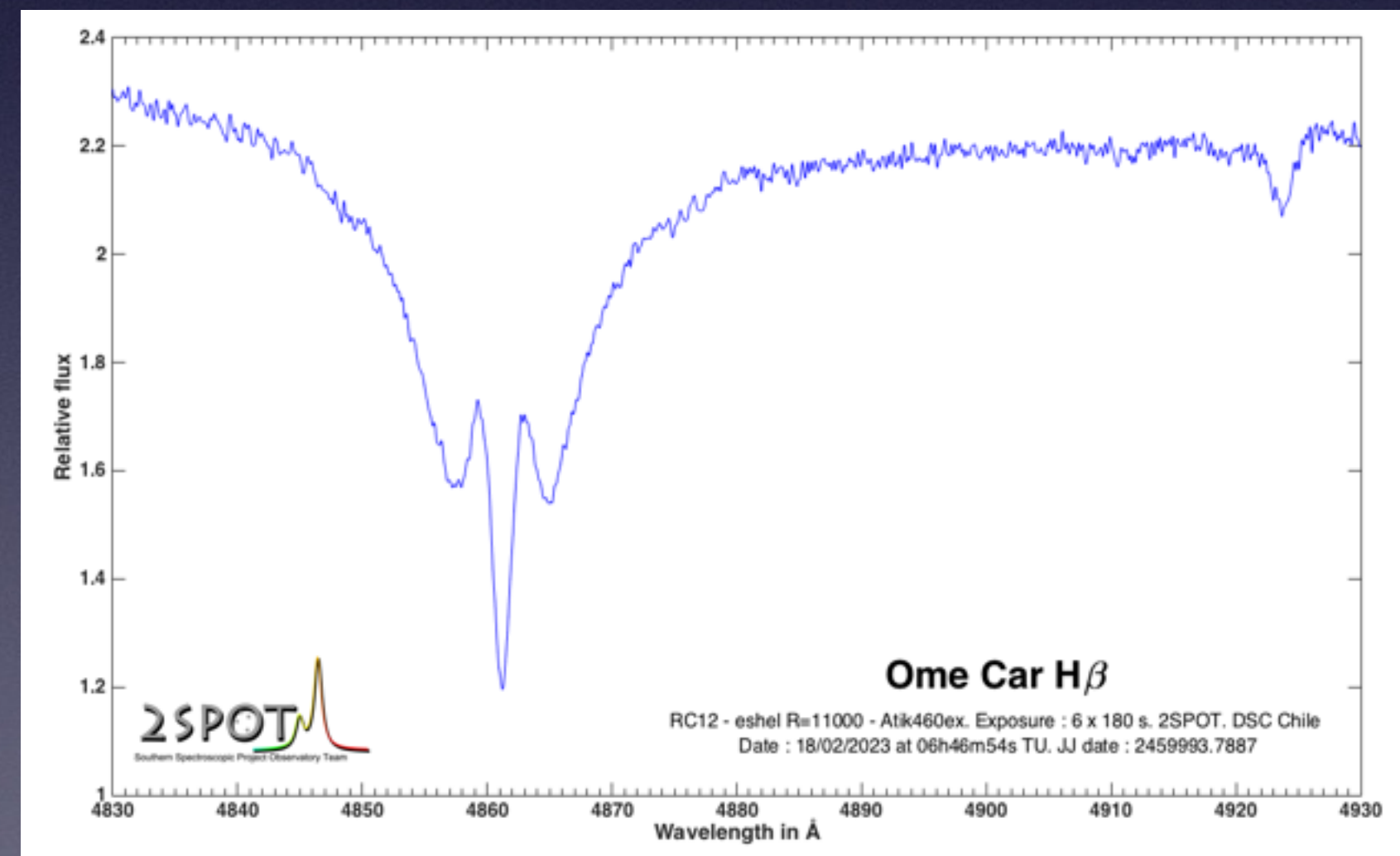
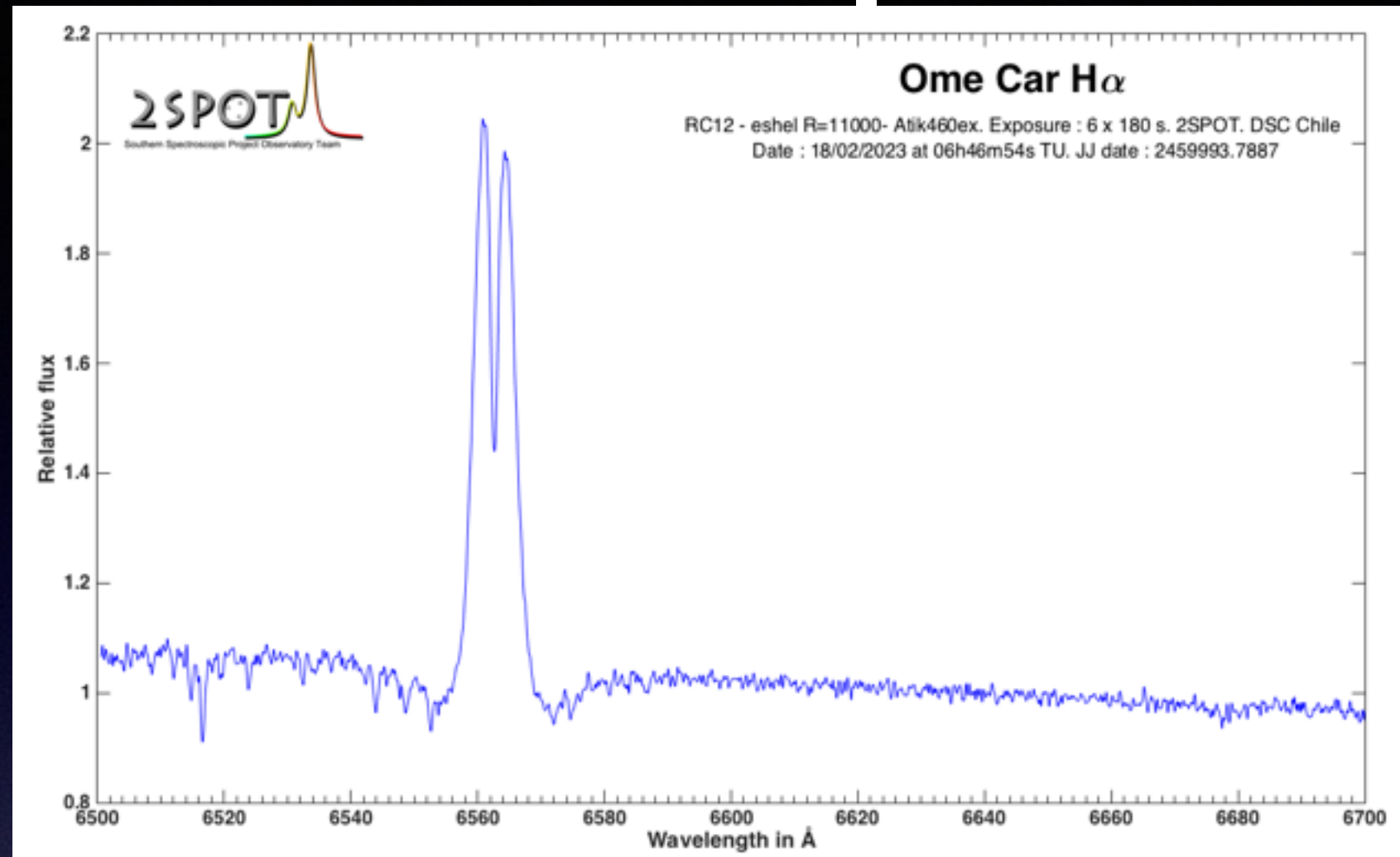
1ère spectre eShel depuis le Chili

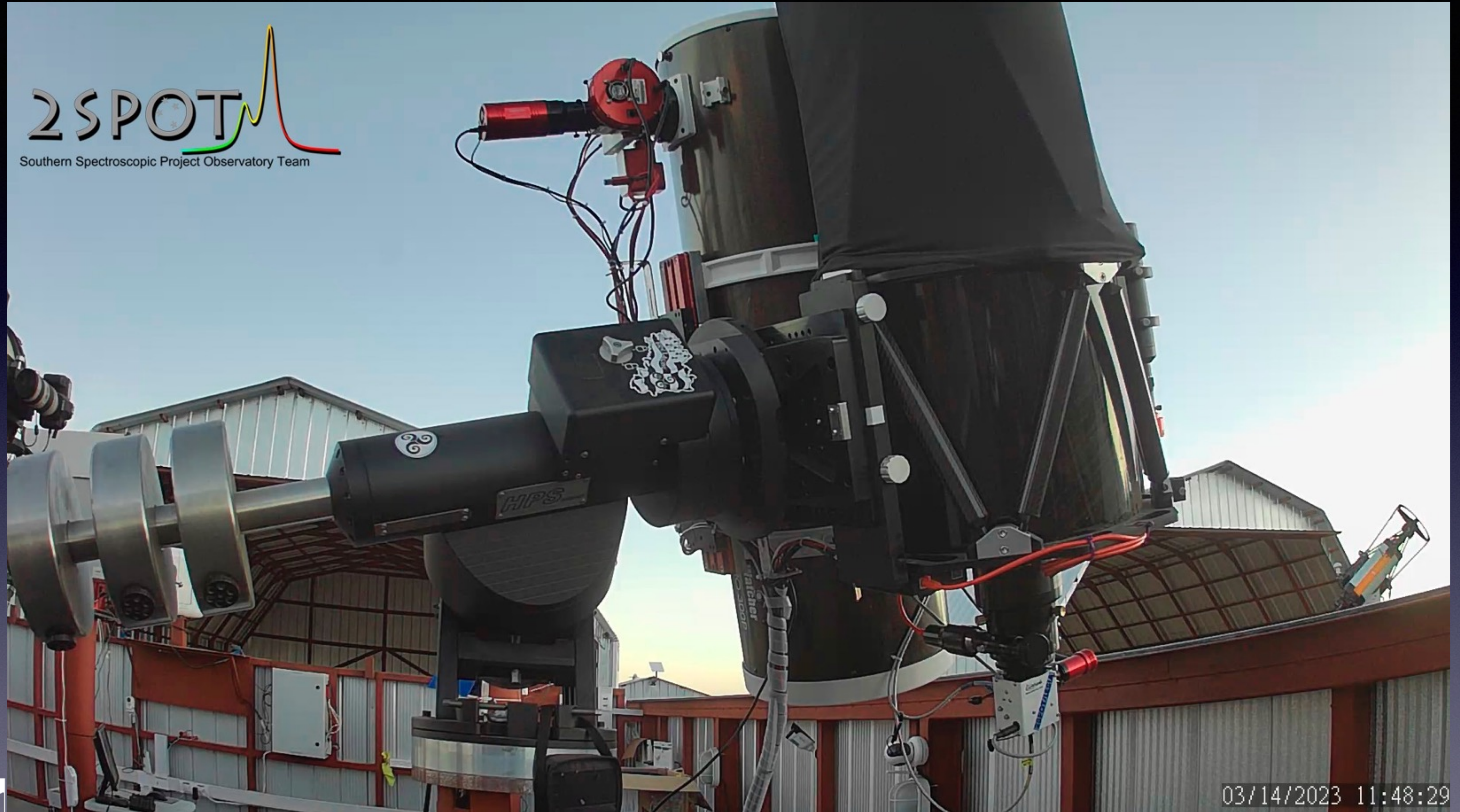
Ome Car
Une étoile Be du Sud



Un exemple de spectre moyenne résolution échelle

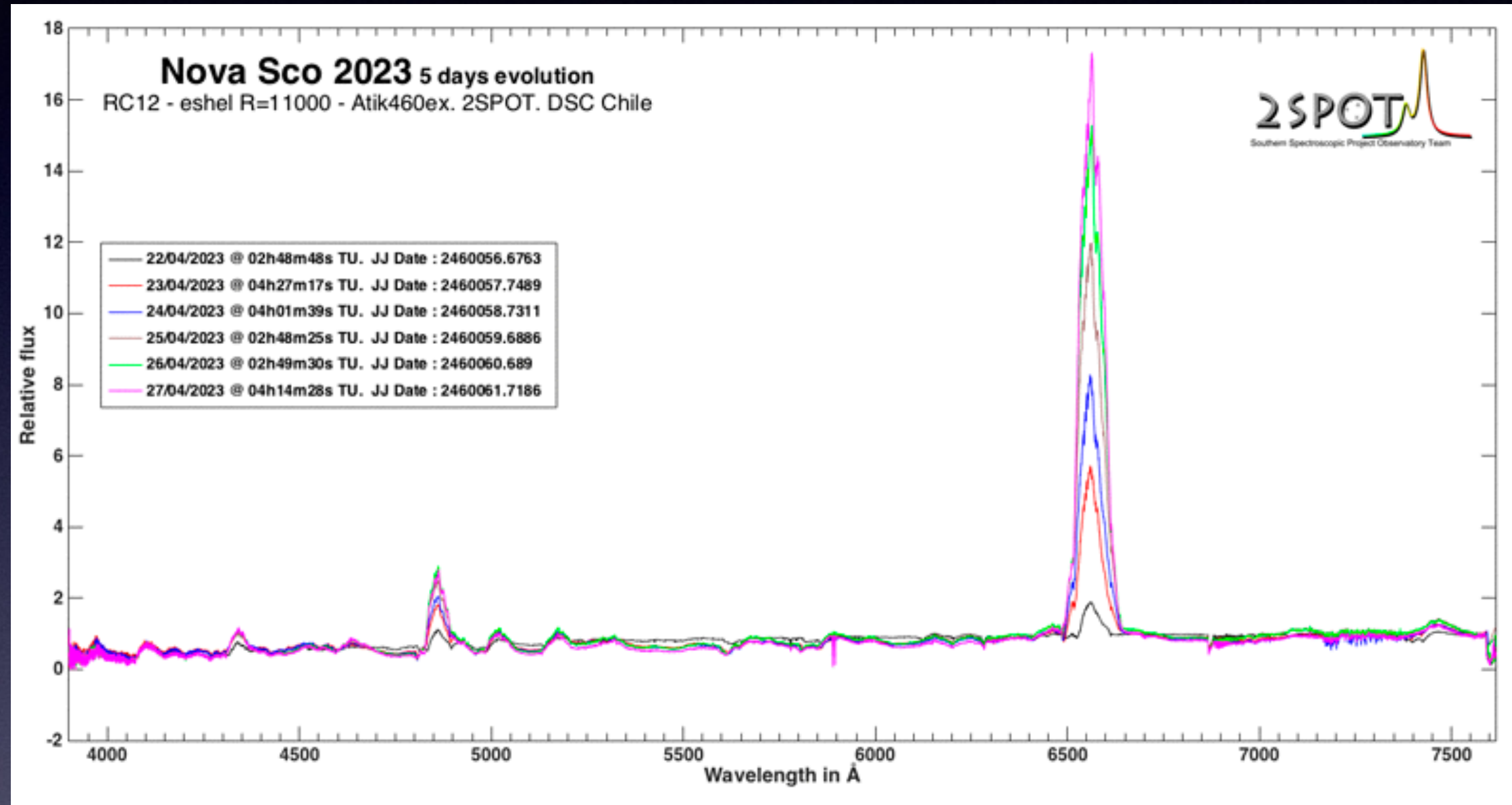
On a réalisé plus de **800** spectres pour la base Bess



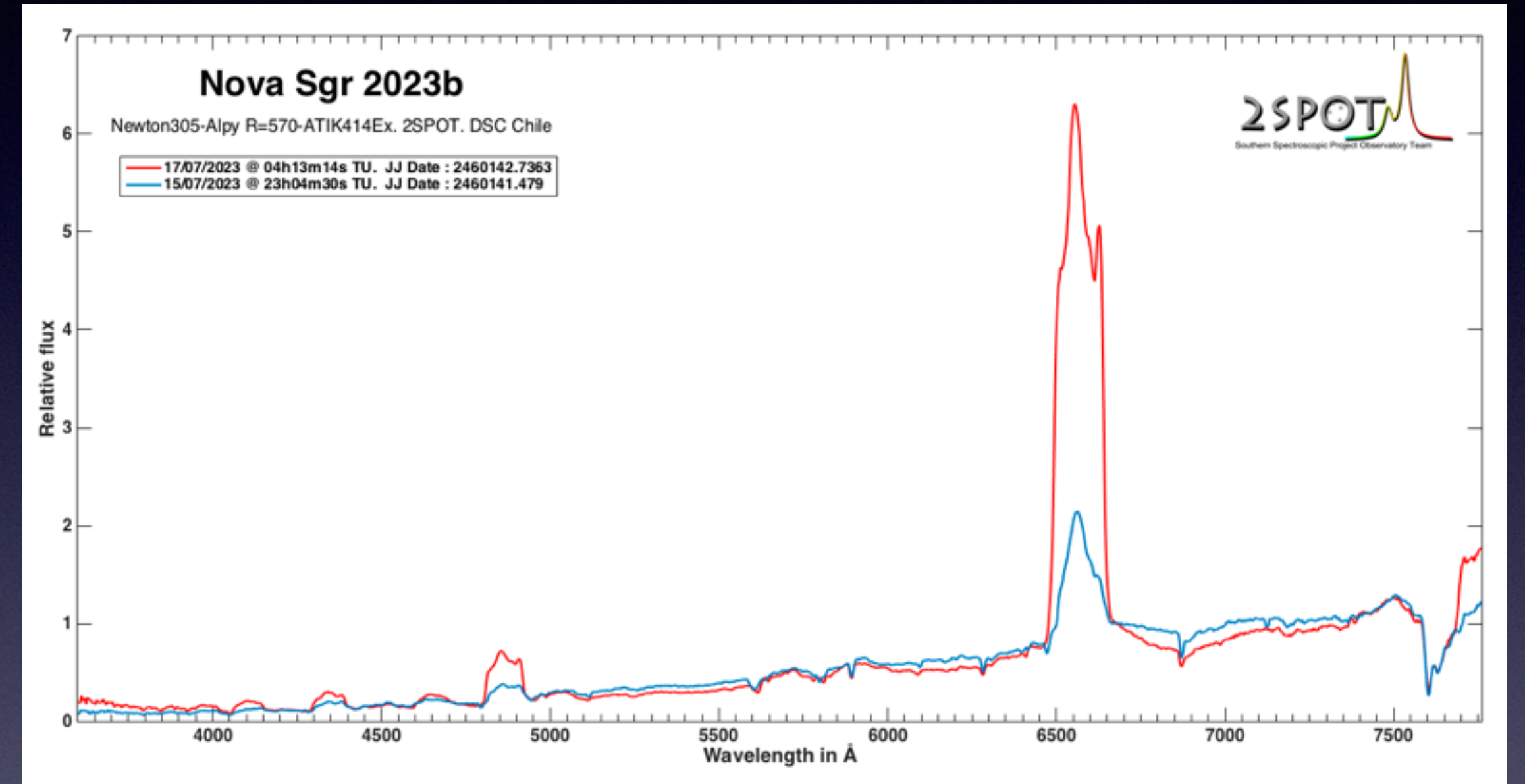


03/14/2023 11:48:29

Suivie de 2 novae en 2023



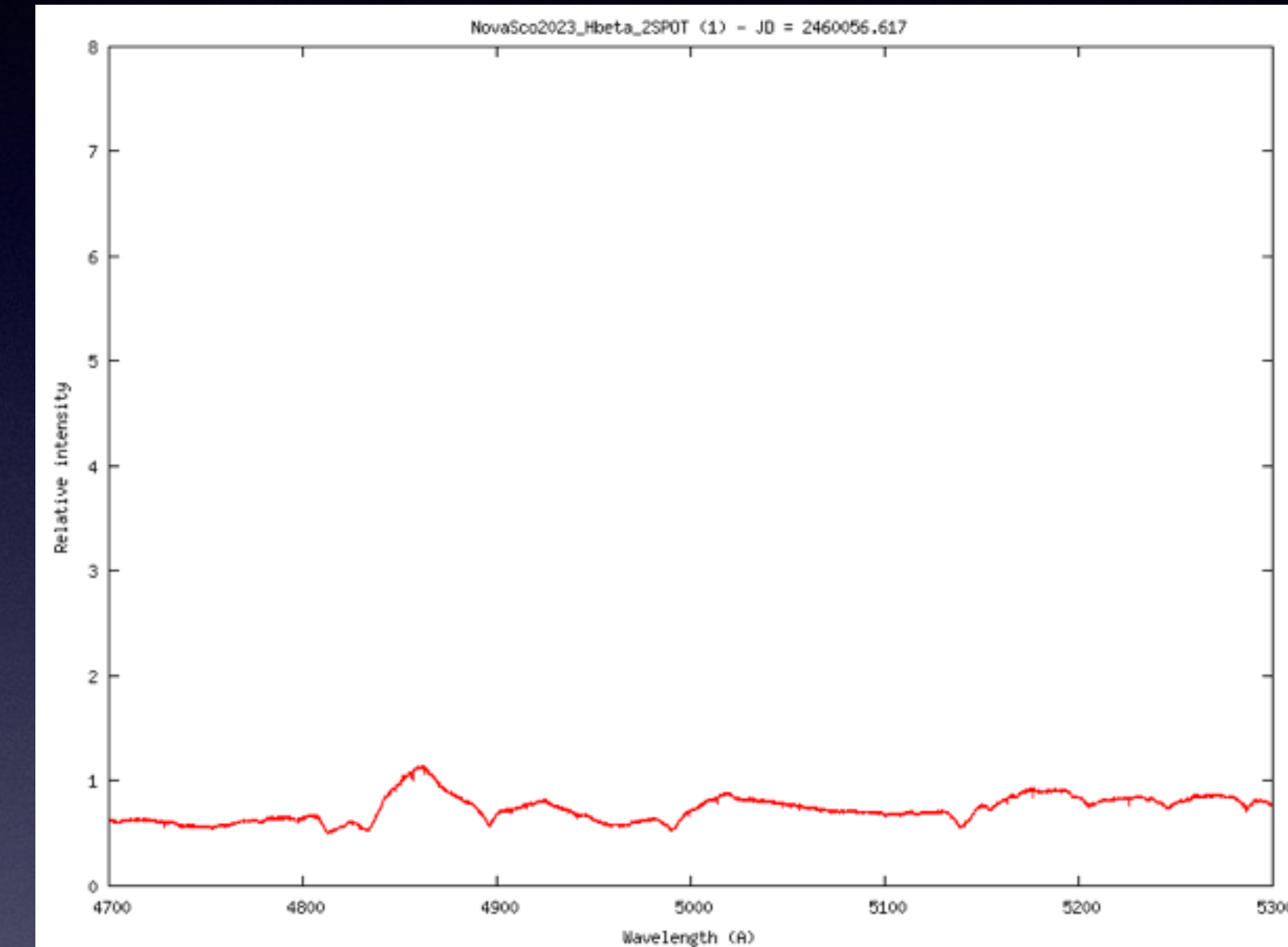
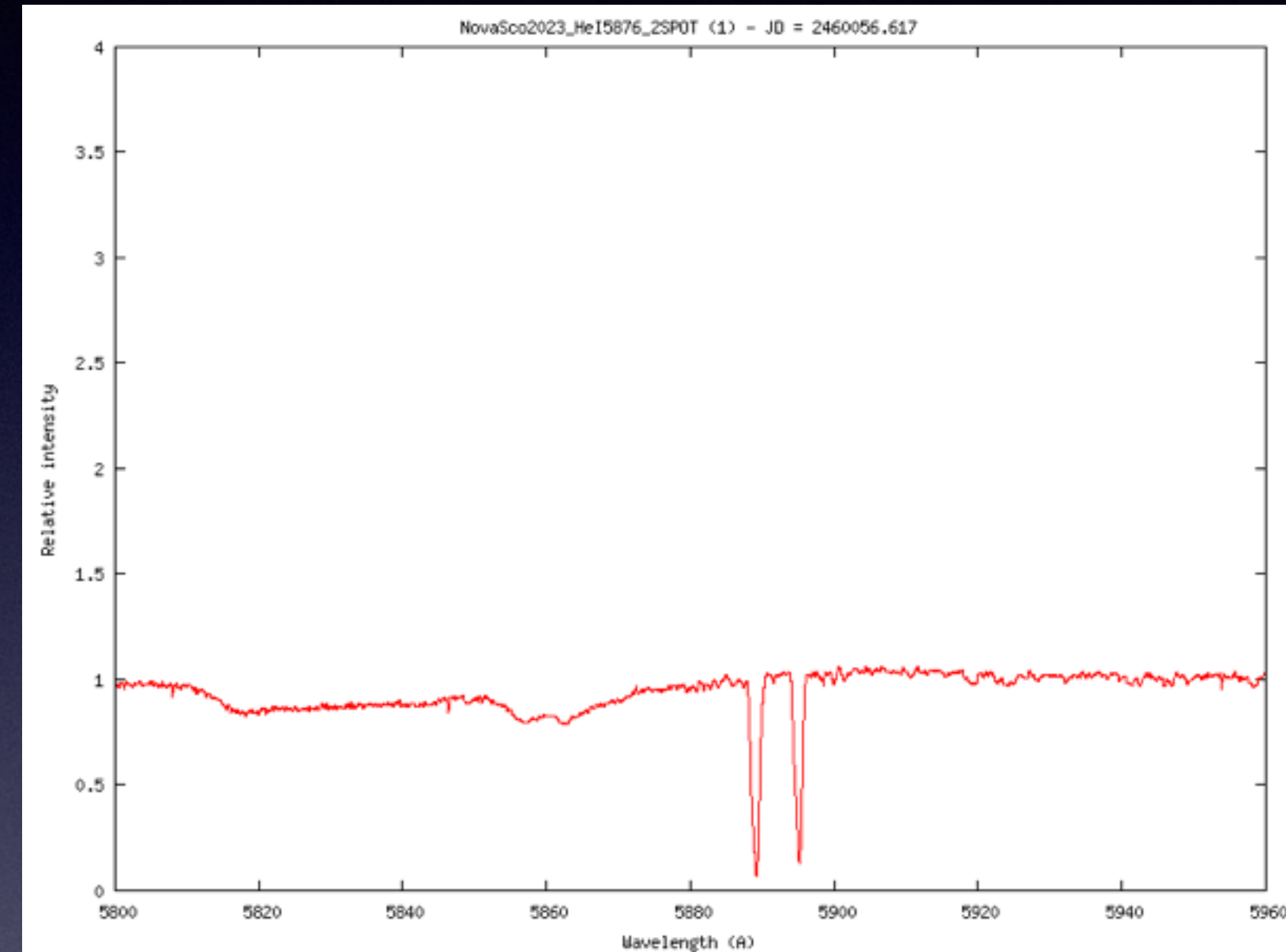
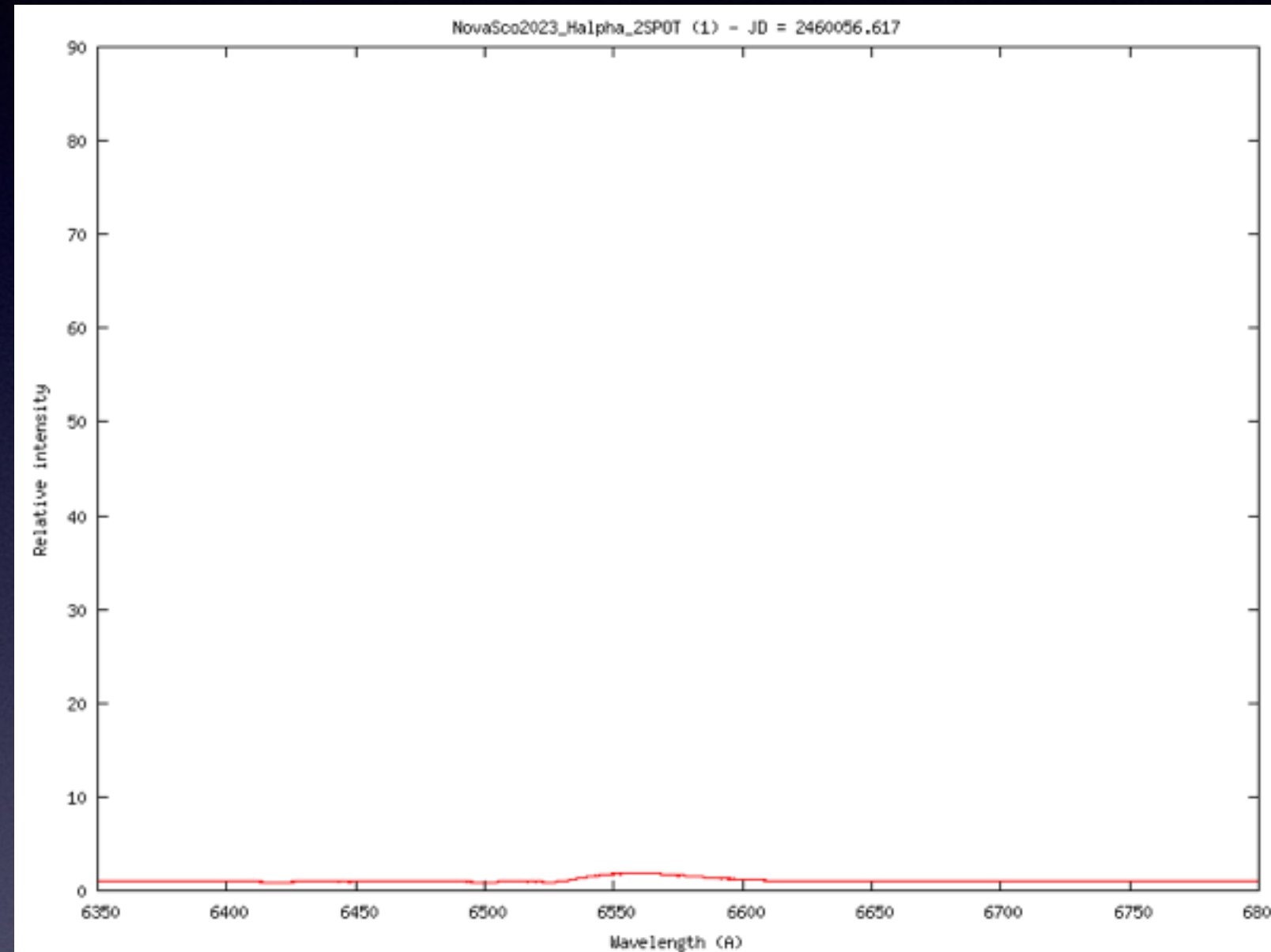
Nova Sco 2023



Nova Sgr 2023b



Suivie de 2 novae



Evolution de Nova Sco 2023 en 1 mois



Rencontres du Ciel et de l'Espace 2024

Publications sur les 2 novae

Circulaire de l'UAI

Nova Sco 2023

ATel #16004 du 22 Avril 2023

ATel #16006 du 23 avril 2023

CBET #5245 du 24 avril 2023

ATel #16036 du 12 mai 2023

Nova Sgr 2023b

ATel #16135 du 23 juillet 2023

CBET#5278 du 17 juillet 2023

O. Garde writes, on behalf of the Southern Spectroscopic Observatory Project Team (2SPOT), which uses a "Deep Sky Chile" telescope at Ruo Hurtado, Chile, that spectroscopy of PNV J17224490-4137160 was obtained on Apr. 22.25 UT with an RC12 reflector (+ sShel echelle fibre spectrograph; resolution 11000) at Deep Sky Chile Observatory, revealing complex lines including Balmer lines with strong absorption (H_alpha velocity -1800 km/s). Na I D shows as a single strong absorption (-1700 km/s) with a broad P-Cyg-like emission, and it indicates a signification interstellar extinction. Iron, Mg, and Ca lines are also visible.

The spectrum appears to be that of a classical nova, and it can be viewed at website URL:

<http://www.spectro-aras.com/forum/viewtopic.php?f=36&t=3090>.

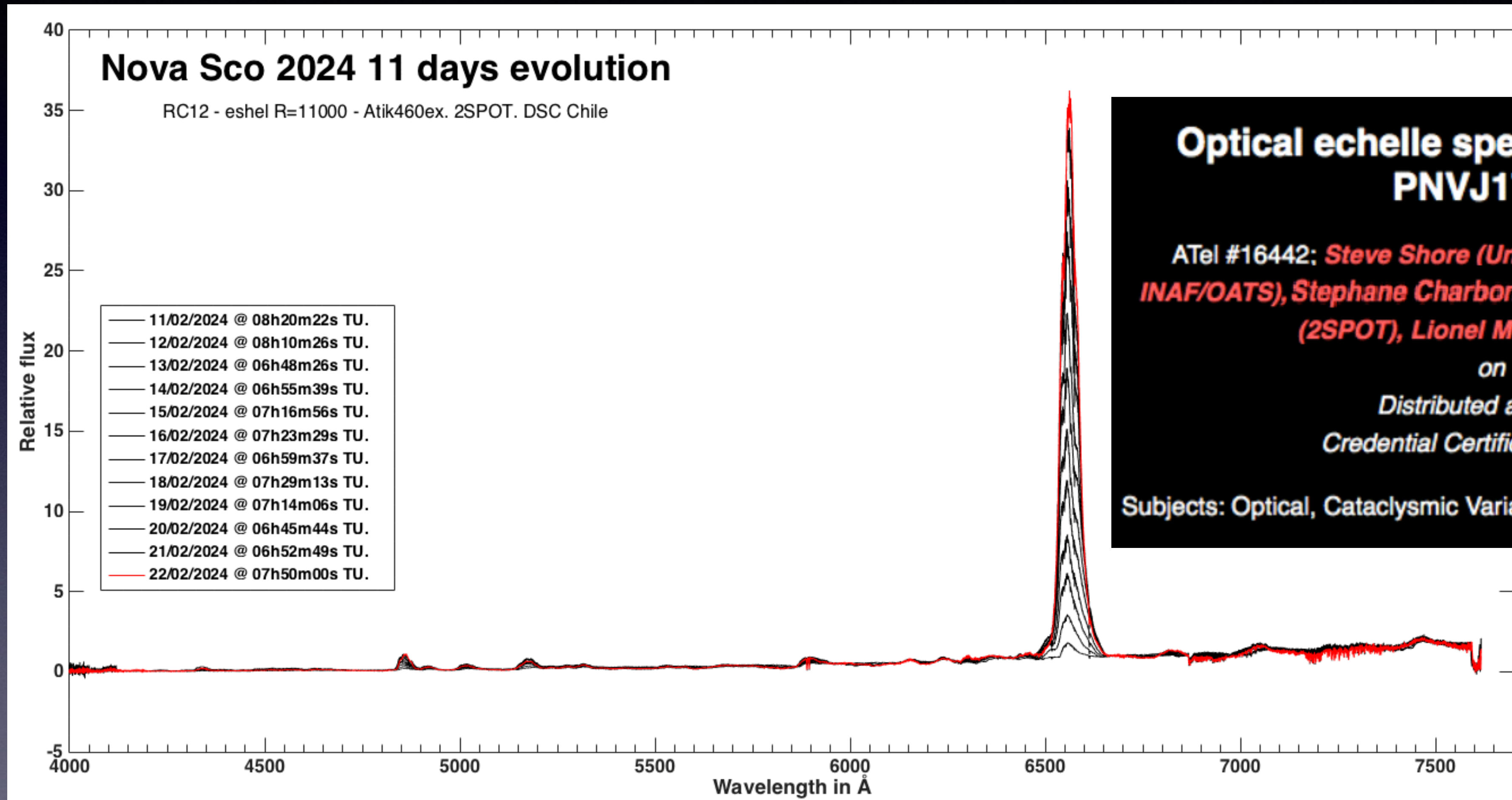
E. Kazarovets writes that the permanent GCVS designation V1716 Sco has been assigned to this nova.

NOTE: These 'Central Bureau Electronic Telegrams' are sometimes superseded by text appearing later in the printed IAU Circulars.



Nova V1723 Sco (Nova Sco 2024)

24 spectres entre le 11 février et 11 mars 2024



Optical echelle spectroscopy of the classical nova PNVJ17261813-133809354

ATel #16442: *Steve Shore (Univ. Pisa, INAF/OATS), Steve Shore (Univ. of Pisa, INAF/OATS), Stephane Charbonnel, (2SPOT), Olivier Garde (2SPOT), Pascal Le Du (2SPOT), Lionel Mulato (2SPOT), Thomas Petit (2SPOT)*

on 13 Feb 2024; 17:24 UT

Distributed as an Instant Email Notice Novae

Credential Certification: S. N. Shore (shore@df.unipi.it)

Subjects: Optical, Cataclysmic Variable, Nova, Transient



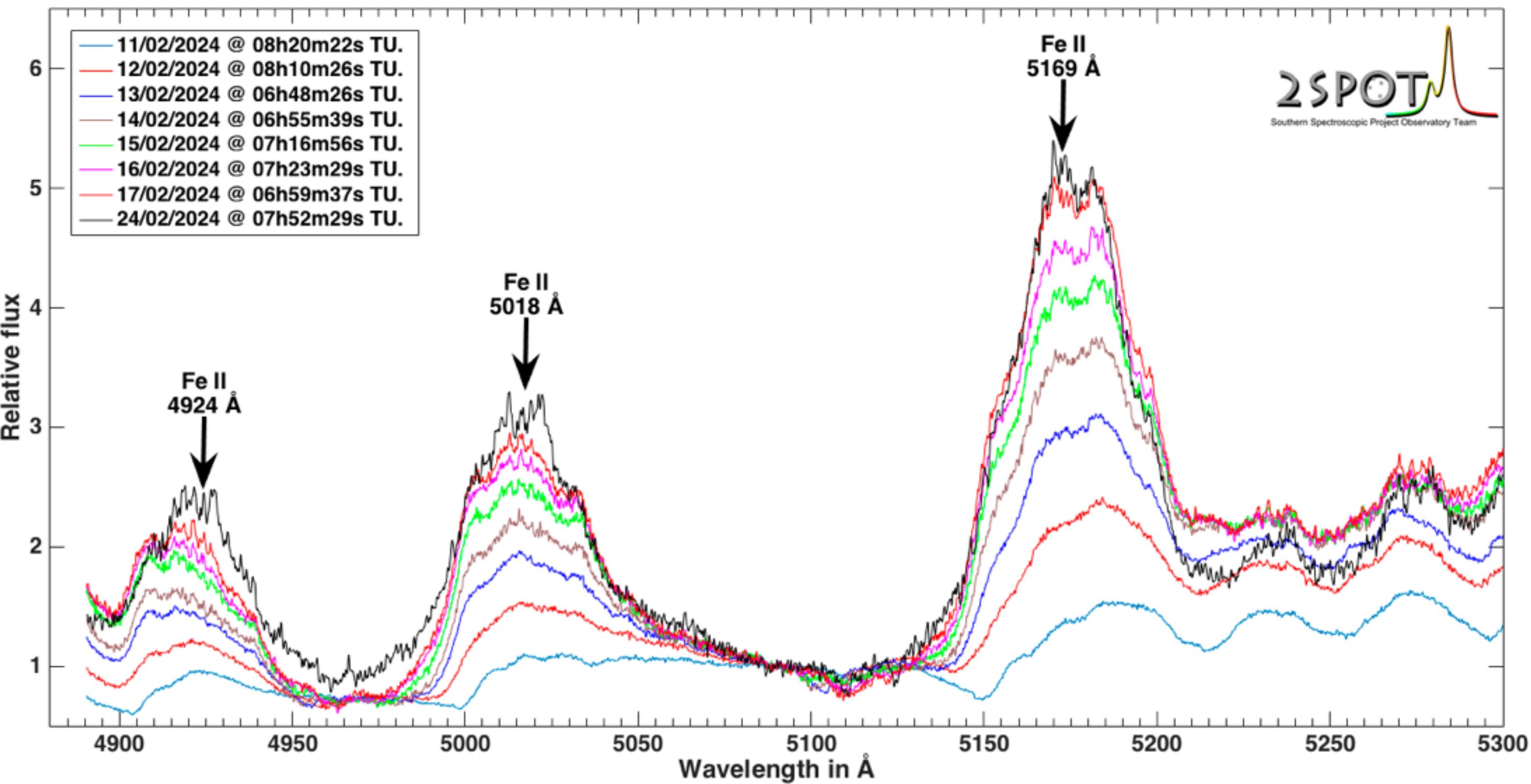
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Nova V1723 Sco (Nova Sco 2024)

Evolutions de la raie H Alpha et du Fer II

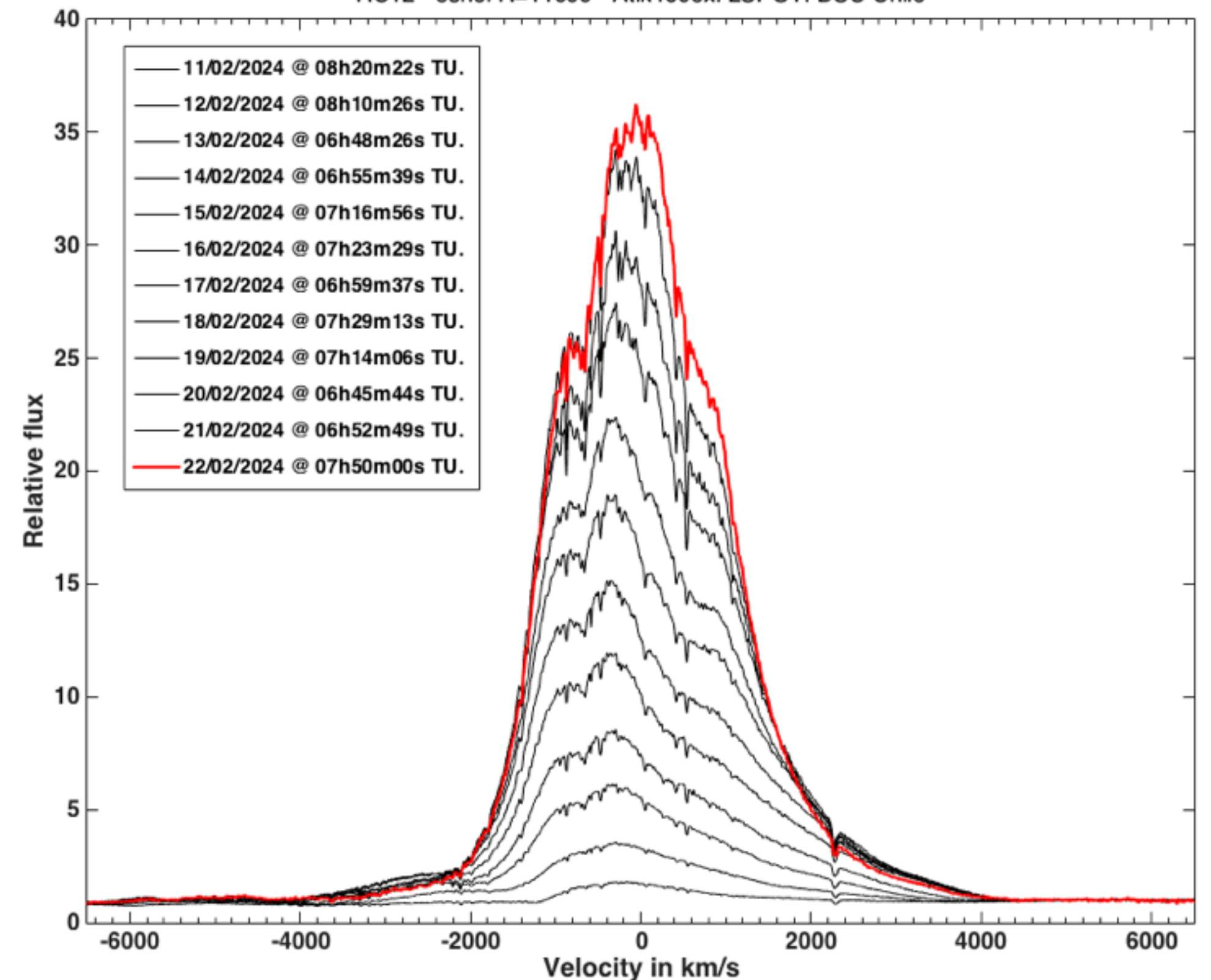
V1723 Sco Fe II lines evolution in 13 days

RC12 - eshel R=11000 - Atik460ex. 2SPOT. DSC Chile



Nova Sco 2024 H α evolution in 11 days

RC12 - eshel R=11000 - Atik460ex. 2SPOT. DSC Chile



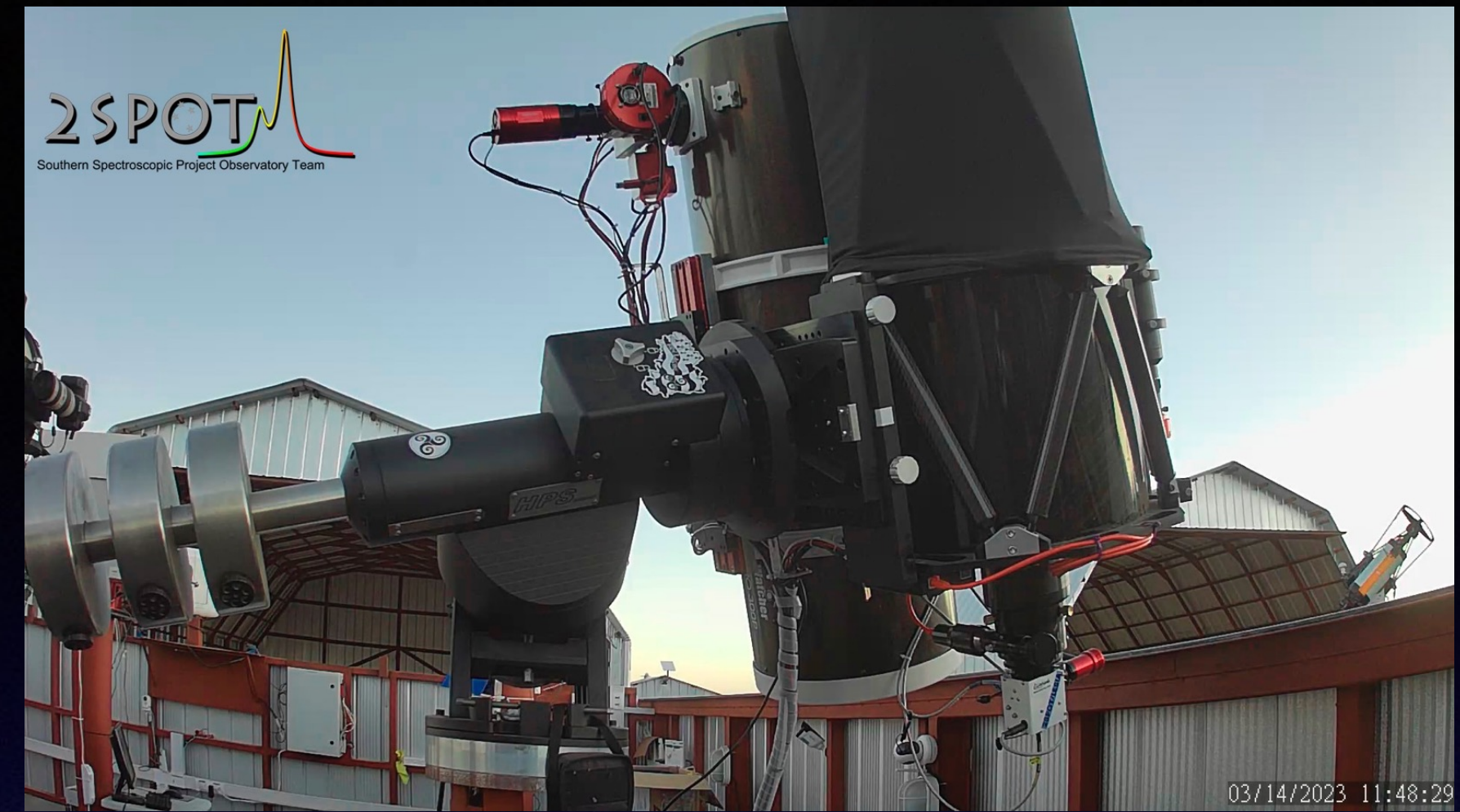
Un article dans la revue l'Astronomie (N°173 Juillet/Aout 2023)



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Notre force

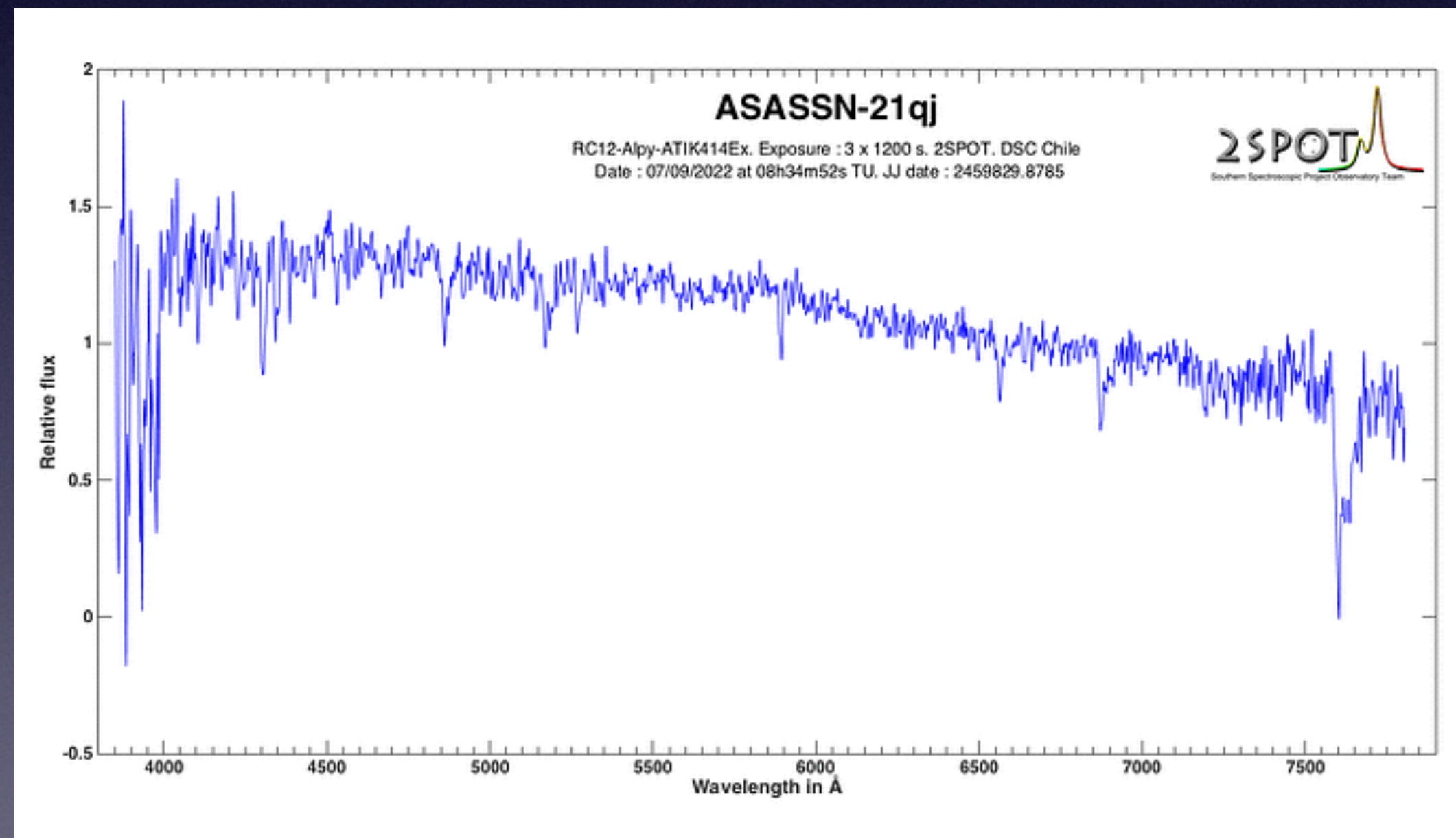


- **Réactivité** pour les observations : on peut réaliser un spectre quelques heures après la découverte en programmant la cible dans notre pipeline.
- **100% process automatique** : ce qui nous permet d'avoir tout les résultats quelques heures après la fin de la nuit.
- Un système polyvalent avec **2** spectrographes complémentaires
- Un spectrographe basse résolution qui nous permet d'obtenir des cibles dont la magnitude est supérieur à 17.



Réactivité

ASASSN-21qj une cible que nous avons pu faire très rapidement afin de prouver que le spectre était de type Solaire et n'avait pas subi d'évolution. Ce spectre a été très utile pour cette publication (un petit élément dans ce vaste puzzle de données)



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Article | Published: 11 October 2023

A planetary collision afterglow and transit of the resultant debris cloud

[Matthew Kenworthy](#) , [Simon Lock](#), [Grant Kennedy](#), [Richelle van Capelleveen](#), [Eric Mamajek](#), [Ludmila Carone](#), [Franz-Josef Hamsch](#), [Joseph Masiero](#), [Amy Mainzer](#), [J. Davy Kirkpatrick](#), [Edward Gomez](#), [Zoë Leinhardt](#), [Jingyao Dou](#), [Pavan Tanna](#), [Arttu Sainio](#), [Hamish Barker](#), [Stéphane Charbonnel](#), [Olivier Garde](#), [Pascal Le Dû](#), [Lionel Mulato](#), [Thomas Petit](#) & [Michael Rizzo Smith](#)

[Nature](#) **622**, 251–254 (2023) | [Cite this article](#)



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- * **The Astrophysical Journal Letters** du 10 juin 2022 : A Speed Bump with SN 2021aefx
- * **MNRAS** du 17 octobre 2022 : WHTZ 1: A high excitation Planetary Nebula
- * **CBET #5245** april 24th 2023 : Nova Sco 2023 = V1716 Sco
- * **MNRAS** may 8th 2023 : V618 Sgr Galactic eclipsing symbiotic nova detected
- * **L'Astronomie** N°173 July/August 2023 : New star in Scorpius
- * **CBET #5278** July, 17th 2023 : TCP J17525020-2024150 = V6598 Sgr
- * **NATURE** october 11th 2023 : A planetary collision afterglow
- * **CBET #5346** du 12 février 2024 : Nova Sco2024 = V1723 Sco
- * **Astronomische Nachrichten du 19 mars 2024** : Gaia23ckh : symbiotic outburst of V390 Sco 17 juillet 2023
- * **L'Astronomie** N°177 : Exoplanetary collision

Spectroscopic classification of ASASSN-24by (AT 2024epj) as a classical nova in the LMC

ATel #16545; **J. Merc (Charles University), T. Love, P. Velez, H. Barker (ARAS Group), S. Charbonnel, O. Garde, P. Le Du, L. Mulato, T. Petit (2SPOT Team)**

on 21 Mar 2024; 08:31 UT

Credential Certification: Jaroslav Merc (jaroslav.merc@mff.cuni.cz)

Subjects: Optical, Nova, Transient

Received January 11, 2024; | Revised March 18, 2024; | Accepted March 19, 2024

DOI: xxx/xxxx

ORIGINAL ARTICLE

Gaia23ckh: Symbiotic outburst of the assumed Mira variable V390 Sco

Jaroslav Merc¹ | Peter Velez² | Stéphane Charbonnel³ | Olivier Garde³ | Pascal Le Du³ | Lionel Mulato³ | Thomas Petit³ | Jan Skowron⁴

Early spectroscopy of the classical LMC nova AT2 024fjh

ATel #16575; **Steven N. Shore (Univ. Pisa; INAF/OATS), Stéphane Charbonnel, Olivier Garde, Pascal Le Du, Lionel Mulato, Thomas Petit (2SPOT, ARAS Group), Hamish Barker, Peter Velez (ARAS Group)**

on 6 Apr 2024; 02:26 UT

Credential Certification: S. N. Shore (shore@df.unipi.it)

Subjects: Nova, Transient



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